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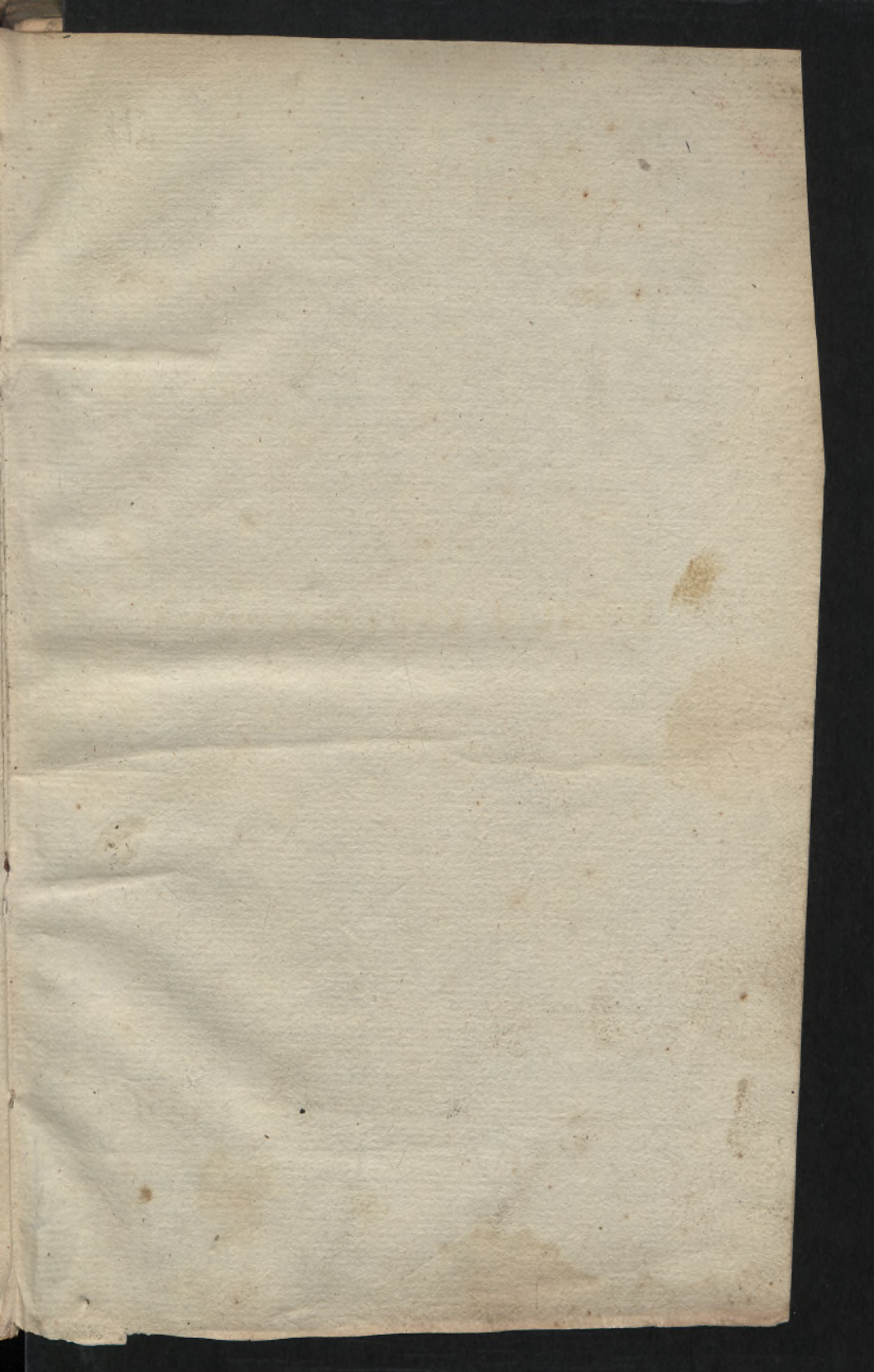


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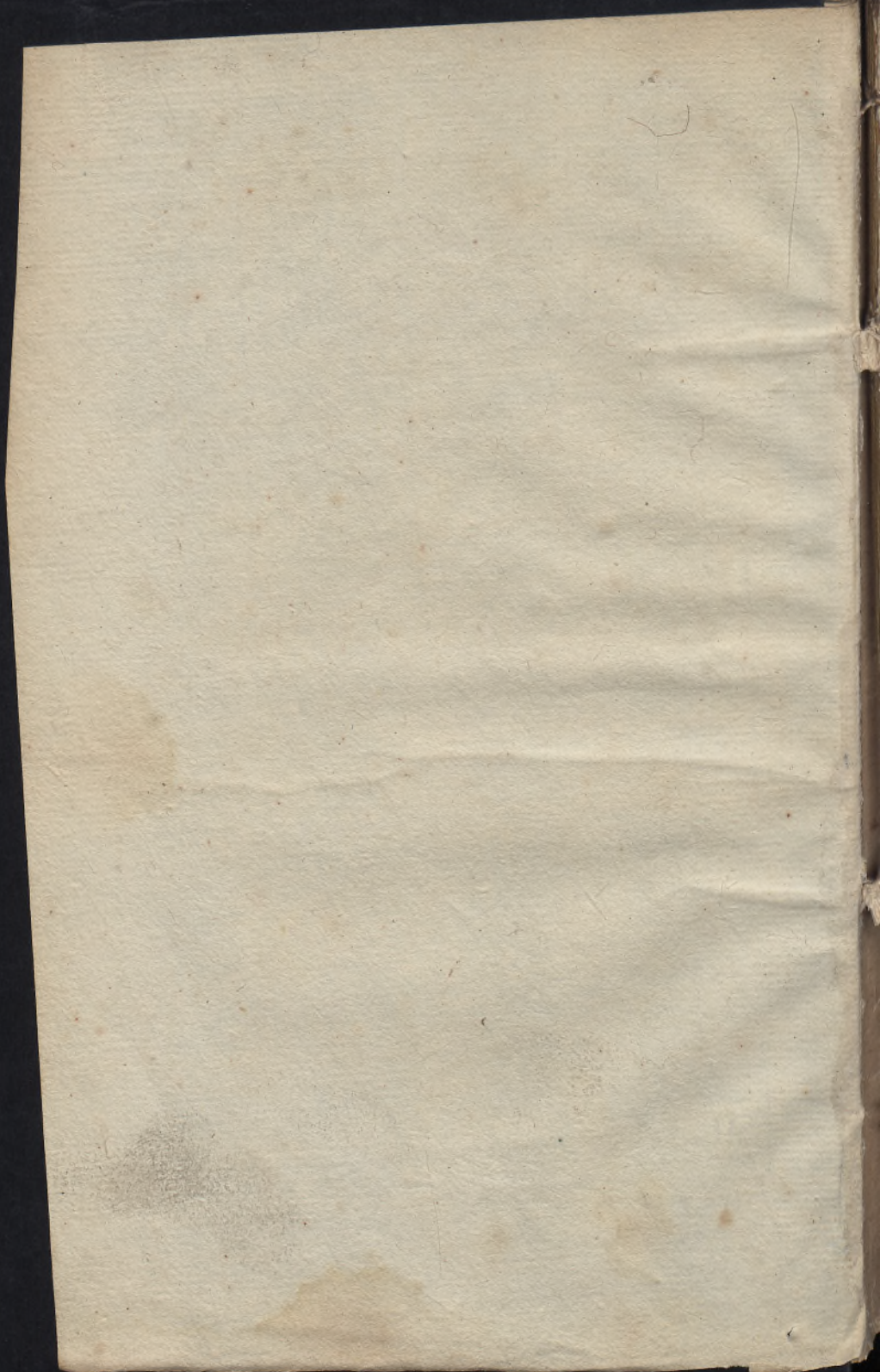
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ASTRONOMY  
OF THE  
SATELLITES

OF THE

*Earth, Jupiter and Saturn:*

Grounded upon Sir *Isaac Newton's* Theory of  
the Earth's SATELLITE.

The THEORY explain'd, and made easy to the  
meanest Capacity, in calculating the true Place of  
the Moon:

And freed from the Errors printed in the said Theory, by  
*Dr. Gregory, Dr. Harris,* and several other Authors.

By which now the Place of the Moon, and Eclipses of the Lumina-  
ries, are found to a very great Exactness.

A L S O

New Tables of the Motions of the Satellites of *Jupi-  
ter and Saturn,* (founded upon the Observations of  
*Mr. Flamsteed, Mr. Cassini, Mr. Hugen, Dr. Halley*  
and *Mr. Pound,*) from the Vernal Equinox:

By which their Places and Positions, in respect of one ano-  
ther, may be exactly determined at any given Time.

*Adapted to the Meridian of London.*

To which is added, A PROBLEM to find the Latitude of the Place by  
the Altitude of the Sun, Moon, or Star, upon any Azimuth; being  
very useful for all Sea-faring Men, as well as Gentlemen and others.

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By *CHARLES LEADBETTER,*  
*Teacher of the MATHEMATICKS.*

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L O N D O N :

Printed for J. WILCOX, at the *Green-Dragon,* in  
*Little-Britain.* M.DCC.XXIX.



YMON  
ADVERTISEMENT.

THE following Mathematical Sciences are taught by the Author hereof, at his House, at *The Hand and Pen* in *Cock-Lane, Shoreditch, London*; or at any Gentleman's Apartment, viz. Vulgar and Decimal Arithmetick, Geometry apply'd to the Mensuration of Superficies and Solids, by Pen and Sliding-Rule; Projection of the Sphere on any Circle; Trigonometry, plain and spherical; Surveying of Land, by any Instrument now in Use; Gauging of all sorts of Vessels, with all the practical Methods used by the Officers of the Excise; Astronomy in all its Branches; Navigation by the Plain and Mercator's Chart, and by the Arch of a great Circle; Geography and the Use of the Globes, with all other Mathematical Instruments whatsoever. Dialling upon any Plane for any Latitude.

Where may be had, *First*, His Treatise of Eclipses for 26 Years. 2. His Sheet of the Conjunction of  $\eta$  4 and  $\delta$  1722, and *Mercury's* Passage over the Sun, *October* 29, 1723. 3. His Astronomical Calendar for 28 Years, ending 1753. 4. His System of the Planets demonstrated. 5. His Sheet of all the Luminarian Eclipses for 35 Years, ending *Anno* 1761. 6. His Compleat System of Astronomy. 7. His Astronomy of the Satellites of the Earth, *Jupiter* and *Saturn*: Also sold by *J. Wilcox*, at the *Green-Dragon* in *Little-Britain, London*.



POLITECHNIKA GDAŃSKA  
Z ZASOBÓW  
POLITECHNIKI GDAŃSKIEJ  
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## R E A D E R,

**I** Here present you with my *Astronomy of the Satellites of the Earth, Jupiter and Saturn. What I mean by the Satellite of the Earth, is the Moon, with Sir Isaac Newton's last Improvement of the Theory of that Planet, which I had published in my Compleat System of Astronomy, had it not been for some very gross Errors printed in the Theory, as laid down by Dr. Gregory, in Pag. 334. of his Latin Astronomy, and carried on in his English Astronomy, Vol. II. pag. 563. by Dr. Harris in his Lexicon Technicum, Vol. I. under the Word Moon: and by the Learned Author of Prælectiones Astronomicæ, pag. 318. and in the English Translation, pag. 345. The most material Faults are the greatest Equation of the Apogee, and the radical Place for the Year 1681. For whoever will be at the pains to examine the Numbers (as I have done) will find that the greatest 66782, and least 43319, Eccentricities, will give the greatest and least Equations of the Lunar Orbit  $7^{\circ} 39' 30''$ ,  $4^{\circ} 57' 56''$ ; and  $12^{\circ} 18' 15''$  for the greatest Equation of the Apogee, which, as they have it, is only  $12^{\circ} 15' 4''$ . By comparing these Numbers with the Times of the visible Eclipses of the Sun in the Years 1715, 1722, 1724, 1726, which were carefully observed at London, I am satisfied that*



the greatest Equation of the Apogee is  $12^{\circ} 18' 15''$  as I have it in these my new Tables, and reduced them to the Meridian of London.

By finding the true Times of the Conjunction, and Opposition of the Sun and Moon by this Theory, it may seem, at first, to them unskill'd in these Matters, to be almost an Impossibility; but after due Consideration it will appear, that those Equations, that depend upon the Distance of the Moon from the Sun, vanish, which are the 5th, 6th and 7th; and the first, second, and third Equations, alter but little, in a small Space of Time; so that regard is chiefly to be had to the fourth Equation, with which work as I have shewed in my Compleat System, Precept 7th. until you find the Orbit-Place of the Moon the same with the Sun's true Place, and then you have the middle Time of the true Conjunction or Opposition in the Moon's Orb, to the greatest Exactness imaginable.

Many ingenious Persons have often wish'd, that Tables of the Motions of the Satellites of Jupiter and Saturn were published, that thereby they might know at any time before-hand, how they wou'd appear when observed: Therefore, for the sake of the diligent Observer, I here publish mine, constructed from the Observations of Mr. Huggens, Mr. Flamsteed, Mr. Cassini, Dr. Halley, and Mr. Pound; which, I dare to say, are the correctest the World ever saw. The Method of finding their Places is plain and easy to be understood by any one, though meanly versed in these things.

I have only one thing more to remind my Reader of, and that is, If he has a mind to find the Times of the Eclipses of Jupiter's Satellites, after the Cassinian Method, he must observe, that the periodical Time of the  
first



first Satellite is nearly  $\frac{1}{247}$  Part of the periodical Time of Jupiter from one Apheleon to another; whence the Equations of the Jovial Orbit being turned into Minutes and Seconds of Time, and adapted to those particular Revolutions of the Satellites, will make good the principal Parts of the Equations of these Satellites. By which Directions and easy Tables, any Observer may truly know their Distances from Jupiter, and distinguish one Satellite from another at any given Time; which will both be pleasant and advantagious to him in rectifying the Longitude of Places at Land, by the Times of their Eclipses.

I remain a Friend to the Astronomical Student,

CHARLES LEADBETTER,

SECT. II.

SECT. III.

THE

The Use of the Tables of the Satellites of Jupiter and Saturn. Showing how to calculate the Distance of Jupiter's Satellites, and to distinguish one from another. A Table of the mean Motion of the Moon in Years. A Table of the mean Motion of the Moon in Months and Days from Pag. 22 to Pag. 27.



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T H E  
C O N T E N T S.

## S E C T. I.

	Pag.
<b>T</b> O calculate the true Place of the Moon	I
A Table of the first Equation of the Moon	3
To find the second Equation of the Moon's Apogee	4
To find the first Equation of the Moon's Node	6
To find the second Equation of the Moon	8
To find the third Equation of the Moon	II
To find the second Equation of the Moon's Apogee	13
To find the present Eccentricity of the Moon	16
To find the Elliptic Equation of the Moon	<i>ibid.</i>
To find the Moon's Variation	18
To find the sixth Equation of the Moon	24
To find the seventh Equation of the Moon	25
To find the Moon's Latitude and Reduction	26
To find the Inclination of the Limit	28
An Example of the Sun and Moon's Place	30
A second Example of the Sun and Moon's Place	31

## S E C T. II.

A Problem of the Sphere.	32
--------------------------	----

## S E C T. III.

The Use of the Tables of the Satellites of <i>Jupiter</i> and <i>Saturn</i>	36
Shewing how to calculate the Distance of <i>Jupiter's</i> Satellites, and to distinguish one from another	41
A Catalogue of Observations.	45
A Table of the mean Motion of the Moon in Years cur.	47
A Table of the mean Motion of the Moon in Months and Days, from Pag. 52, to Pag	54
The	



The mean Motion of the Moon in Hours, &c.	65
A Table of the second Equation of the Node and Inclination of the Limit above $4^{\circ} 59' 35''$	67
A Table of the Moon's simple Latitude	68
A Table of the Moon's Reduction and Excess	69
A Table of the hourly Motions, Semidiameters and Horizontal Parallaxes	70, 71
A Table of the Motion of the first Satellite of <i>Jupiter</i>	72
A Table of the Motion of the second Satellite	74
A Table of the Motion of the third Satellite	76
A Table of the Motion of the fourth Satellite	78
A Table of the Distances of the Satellites of <i>Jupiter</i> , from his Body	80
A Table of the Distances of the Satellites of <i>Jupiter</i> , from his Body, &c.	81
A Table of the Motion of the first Satellite of <i>Saturn</i>	82
A Table of the Motion of the second Satellite of $\text{h}$	84
A Table of the Motion of the third Satellite of $\text{h}$	86
A Table of the Motion of the fourth Satellite of $\text{h}$	88
A Table of the Motion of the fifth Satellite of $\text{h}$	90
A Table of the Distances of <i>Saturn's</i> Satellites, from his Body	92
The Table continued	93
A Table of the Number of Days from the first of <i>January</i> to any Day in the Year	94
To find the Horizontal Parallax and Appar. Semidiameter of the Moon, according to the Theory	95

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BOOKS



BOOKS just publish'd, and sold by J. Wilcox in Little Britain.

I. A Compleat System of Astronomy, in two Volumes; containing the Description and Use of the *Sector*; the Laws of Spherick Geometry; the Projection of the Sphere Orthographically and Stereographically upon the Planes of the Meridian, Ecliptic and Horizon; the Doctrine of the Sphere; and the Eclipses of the Sun and Moon for thirty-seven Years: Together with all the Precepts of Calculation. Also, new Tables of the Motions of the Planets, fix'd Stars, and the first Satellite of *Jupiter*; of Right and Oblique Ascensions, and of Logistical Logarithms. To the whole are prefix'd, *Astronomical Definitions*, for the Benefit of young Students. By Charles Leadbetter, Teacher of the Mathematicks. Price 12 s.

II. Astronomy, or, the true System of the Planets demonstrated: Wherein is shewn, by Instrument, their Anomalies, Heliocentric and Geocentric Places, both in Longitude and Latitude, their Aphelions, Perihelions, Retrogradation, Elongation, Parallaxes and Distances from the Sun and Earth, with the Method of computing the Times when *Venus* and *Mercury* may be seen in the Sun's Disk. Also the Moon's Phases and Eclipses of the Luminaries, for any Time past, present, or to come; with proper Cuts to each Planet, by which any Person may in a few Hours, and with great Ease, attain to a perfect Knowledge of the Planetary or Solar System. By Charles Leadbetter. Price 5 s.

III. A new Treatise of the Construction and Use of the *Sector*; containing the Solutions of the principal Problems, by that admirable Instrument in the chief Branches of Mathematicks, viz. Arithmetick, Mensuration, Plain Trigonometry, Projection of the Sphere, Geography, Astronomy, Dialling, &c. illustrated with Variety of necessary Observations and pleasant Conclusions; containing several Applications intirely new. By the late Mr. Samuel Cunn, now carefully revised by Edmund Stone, F. R. S. Price 4 s.

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VI. The Modern Navigator's Complete Tutor, or a Treatise of the whole Art of Navigation, in its Theory and Practice, Curiosity and Utility. By Joshua Kelly. Price 4 s.



## SECTION I.

To Calculate the true Place of the Moon, according to Sir ISAAC NEWTON's last Improvement of his Theory.

1. FROM the Tables of the Sun's Motions in my complete System of Astronomy, calculate the true Place of the Sun, its Distance from the Earth, and Parts answering the Logarithm, and subtract the Mean Anomaly from the Mean Longitude, and you will have the Place of the Sun's Apogæum, as is shewed in the Examples for the first and second Days of January at Noon, equal Time, Anno 1729.

2. Out of these Tables of the Moon's Mean Motions, take out the Longitude, Apogee and Node, to the Equal Time of the Question proposed, and gather the Mean Longitude and Apogee into two distinct Sums; but the Mean Motion of the Node being Retrograde, you must subtract the Motions for Months and Days from the Radical Place, as is usually done, and as you see in the Examples following.

3. Before you can find any of the Moon's Equations, you must observe that these are standing Numbers, viz.

The Earth's	{	Eccentricity	--	1692
	{	Greatest Equation	1°	56' 20"
The greatest Annual Equation	D	Longitude	11'	49"
	D	Apogee	20	0
	D	Node	--	9 30

And these four are always proportional to each other.

B

And



4. To find the first Equation of the Moon's Longitude. The Sun's Equation for the Time of the Question is 28' 6", then by the Logistical Logarithms you must always say,

As 1° 56' 20" the Sun's greatest Equation.	} 116' 20" LL. 2875	}	+
To its present Equation, 28 6			
So D Annual greatest Equation, 11 49			7057
<hr/>			
To D first Equation in Longitude	2 51		13226

Note, this Equation must always be added to the Mean Longitude of the Moon, when the Sun's Mean Anomaly is 0, 1, 2, 3, 4, 5 Signs; but must be subtracted, if the Sun's Mean Anomaly be 6, 7, 8, 9, 10, 11 Signs: from which Theory I have fram'd the following Table.

Enter the Table with the Sign on the Head and Degree of the Sun's Mean Anomaly on the Left hand descending, but with the Degree on the Right hand ascending, if the Sign fall at bottom of the Table, and in the place of meeting, you will have the first, or Annual Equation of the Moon's Longitude; which in this Example (as above) you will find to be 2' 51" to be subtracted. See Gregory's Astr. p. 544.



A Table of the first Equation of the Moon.

Anom. $\odot$	Sign 0 Add.		Sign 1 Add.		Sign 2 Add.		Sign 3 Add.		Sign 4 Add.		Sign 5 Add.		Anom. $\odot$
	1	"	1	"	1	"	1	"	1	"	1	"	
0	0	0	5	47	10	7	11	49	10	21	6	0	30
1	0	12	5	58	10	14	11	49	10	15	5	49	29
2	0	24	6	9	10	21	11	49	10	7	5	38	28
3	0	36	6	19	10	26	11	48	10	1	5	27	27
4	0	48	6	29	10	31	11	48	9	55	5	16	26
5	1	0	6	39	10	37	11	47	9	47	5	5	25
6	1	12	6	49	10	42	11	47	9	40	4	53	24
7	1	24	6	59	10	47	11	45	9	33	4	42	23
8	1	36	7	9	10	52	11	44	9	26	4	30	22
9	1	48	7	19	10	56	11	42	9	19	4	18	21
10	1	59	7	28	11	1	11	41	9	11	4	6	20
11	2	11	7	37	11	5	11	39	9	2	3	54	19
12	2	23	7	46	11	9	11	36	8	54	3	43	18
13	2	35	7	55	11	14	11	34	8	46	3	31	17
14	2	46	8	4	11	19	11	31	8	37	3	19	16
15	2	58	8	13	11	22	11	29	8	28	3	7	15
16	3	10	8	22	11	25	11	25	8	20	2	54	14
17	3	22	8	31	11	28	11	22	8	11	2	42	13
18	3	34	8	39	11	30	11	19	8	1	2	30	12
19	3	46	8	47	11	32	11	15	7	52	2	18	11
20	3	57	8	55	11	35	11	11	7	43	2	5	10
21	4	9	9	3	11	38	11	6	7	34	1	52	9
22	4	20	9	11	11	40	11	2	7	24	1	40	8
23	4	32	9	19	11	42	10	58	7	14	1	28	7
24	4	43	9	27	11	43	10	53	7	3	1	15	6
25	4	54	9	34	11	45	10	48	6	53	1	2	5
26	5	4	9	41	11	46	10	43	6	43	0	49	4
27	5	15	9	47	11	47	10	38	6	33	0	37	3
28	5	27	9	54	11	48	10	32	6	22	0	25	2
29	5	37	10	1	11	48	10	26	6	11	0	12	1
30	5	47	10	7	11	49	10	21	6	0	0	0	0
	Sign 11 Sub.	Sign 10 Sub.	Sign 9 Sub.	Sign 8 Sub.	Sign 7 Sub.	Sign 6 Sub.							



5. To find the Annual or first Equation of the Moon's Apogee.

BY remembering what the Sun's Equation is for the present Time, you must say,

As $1^{\circ} 56' 20''$ the greatest Equation, $116' 20''$ LL	2875	} +
To its present Equation. -- --	28 6	
So greatest Annual Equation Apog. 20 0	4771	
To present Equation Apogee.	4 50	10940

This Equation is always to be added to the Mean Place of the Moon's Apogee; if the Sun's Mean Anomaly be 6, 7, 8, 9, 10, 11 Signs; but subtracted when the Sun's Mean Anomaly is 0, 1, 2, 3, 4, 5 Signs, as the following Table sheweth, which I have made from the Theory.

Enter the following Table with the Sun's Mean Anomaly, as directed in the 4th, and you have the Equation answering, to be added to, or subtracted from the Mean Place of the Moon's Apogee; and you will have it Equated the first time.

0	0	0	0	0	0	0	0	0	0
1	0	0	0	0	0	0	0	0	0
2	0	0	0	0	0	0	0	0	0
3	0	0	0	0	0	0	0	0	0
4	0	0	0	0	0	0	0	0	0
5	0	0	0	0	0	0	0	0	0
6	0	0	0	0	0	0	0	0	0
7	0	0	0	0	0	0	0	0	0
8	0	0	0	0	0	0	0	0	0
9	0	0	0	0	0	0	0	0	0
10	0	0	0	0	0	0	0	0	0
11	0	0	0	0	0	0	0	0	0
12	0	0	0	0	0	0	0	0	0

A



A Table of the first Equation of the Moon's Apogee.

Anom. $\odot$	Sign 0 Sub.	Sign 1 Sub.	Sign 2 Sub.	Sign 3 Sub.	Sign 4 Sub.	Sign 5 Sub.	Anom. $\odot$
0	0	9	49	17	8	20	0
1	0	20	10	7	17	19	20
2	0	41	10	24	17	29	20
3	1	1	10	42	17	39	19
4	1	22	10	59	17	48	19
5	1	42	11	16	17	58	19
6	2	3	11	33	18	7	19
7	2	23	11	49	18	16	19
8	2	43	12	6	18	24	19
9	3	4	12	22	18	32	19
10	3	24	12	39	18	39	19
11	3	44	12	55	18	46	19
12	4	4	13	10	18	53	19
13	4	24	13	26	19	0	19
14	4	44	13	41	19	6	19
15	5	4	13	56	19	12	19
16	5	24	14	10	19	18	19
17	5	44	14	25	19	23	19
18	6	3	14	39	19	28	19
19	6	23	14	53	19	33	19
20	6	43	15	7	19	37	18
21	7	2	15	21	19	41	18
22	7	21	15	33	19	44	18
23	7	40	15	46	19	48	18
24	7	59	15	58	19	51	18
25	8	18	16	11	19	53	18
26	8	36	16	23	19	55	18
27	8	54	16	35	19	57	17
28	9	13	16	46	19	58	17
29	9	31	16	57	19	59	17
30	9	49	17	8	20	0	17
	Sign 11 Add.	Sign 10 Add.	Sign 9 Add.	Sign 8 Add.	Sign 7 Add.	Sign 6 Add.	



6. To find the Annual or first Equation of the Moon's Node.

Here you must take the present Equation of the Sun and by the Logistical Logarithms in my Astronomy, say,

As 1° 56' 20" the Sun's greatest Equation,	}	116' 20" LL 2875	}		
To its present Equation,				28 6	3294
So greatest Equation of the Node,				9 30	8004

To the present Equat. of the Node, 2 18. 14173

This Equation is always to be added to the Mean Place of the Node, if the Sun's Mean Anomaly be 0, 1, 2, 3, 4, 5 Signs, but subtracted if it be 6, 7, 8, 9, 10, 11 Signs, the Sum or Difference is the Place of the Node the first time Equated.

0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30
0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1	1
2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2	2
3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3	3
4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4	4
5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5	5
6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6	6
7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7	7
8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8
9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9	9
10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10	10
11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11	11
12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12	12
13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13	13
14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14	14
15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15	15
16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16	16
17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17	17
18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18	18
19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19	19
20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20	20
21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21	21
22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22	22
23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23	23
24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24	24
25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25	25
26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26	26
27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27	27
28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28	28
29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29	29
30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30	30

A



A Table of the Annual, or first Equation of the Moon's Node.

Anom. $\odot$	Sign 0 Add.		Sign 1 Add.		Sign 2 Add.		Sign 3 Add.		Sign 4 Add.		Sign 5 Add.		Anom. $\odot$
	1	11	1	11	1	11	1	11	1	11	1	11	
0	0	0	4	40	8	9	9	30	8	19	4	50	30
1	0	9	4	48	8	14	9	30	8	14	4	41	29
2	0	19	4	57	8	19	9	30	8	9	4	32	28
3	0	29	5	5	8	24	9	30	8	4	4	23	27
4	0	39	5	13	8	28	9	29	7	58	4	14	26
5	0	48	5	21	8	32	9	29	7	53	4	5	25
6	0	58	5	29	8	36	9	28	7	48	3	56	24
7	1	8	5	37	8	40	9	27	7	42	3	47	23
8	1	18	5	45	8	44	9	26	7	35	3	38	22
9	1	27	5	53	8	48	9	25	7	29	3	29	21
10	1	37	6	0	8	52	9	24	7	23	3	19	20
11	1	46	6	8	8	55	9	22	7	16	3	9	19
12	1	56	6	15	8	58	9	20	7	9	2	59	18
13	2	5	6	22	9	2	9	18	7	3	2	50	17
14	2	15	6	30	9	5	9	16	6	56	2	40	16
15	2	24	6	37	9	7	9	14	6	49	2	30	15
16	2	34	6	44	9	10	9	12	6	42	2	20	14
17	2	43	6	50	9	12	9	9	6	35	2	10	13
18	2	53	6	57	9	15	9	6	6	27	2	0	12
19	3	2	7	4	9	17	9	3	6	20	1	50	11
20	3	11	7	11	9	19	8	59	6	12	1	40	10
21	3	20	7	17	9	21	8	56	6	4	1	30	9
22	3	29	7	23	9	22	8	52	5	56	1	20	8
23	3	38	7	29	9	24	8	49	5	49	1	10	7
24	3	47	7	35	9	25	8	45	5	41	1	0	6
25	3	56	7	41	9	26	8	41	5	33	0	50	5
26	4	5	7	46	9	27	8	37	5	24	0	40	4
27	4	14	7	52	9	28	8	33	5	16	0	30	3
28	4	23	7	58	9	29	8	28	5	8	0	20	2
29	4	31	8	4	9	29	8	23	4	59	0	10	1
30	4	40	8	9	9	30	8	19	4	50	0	0	0
	Sign 11 Sub.		Sign 10 Sub.		Sign 9 Sub.		Sign 8 Sub.		Sign 7 Sub.		Sign 6 Sub.		



7. To find the Second Equation of the Moon

**T**His Equation is greatest when the Sun is in the  
 Occants, or  $45^{\circ}$  distant from the Moon's Apogee, and  
 is then  $3' 56''$  if the Sun be in Perigeon; but only  $3' 34''$   
 in Apogeeum, and at a mean Distance from the Earth  $3' 45''$ .  
 Therefore, subtract the Place of the Moon's Apogee the  
 first time Equated, from the Sun's true Place, double the  
 Remainder and say, As Radius, to Sine of that double  
 Distance of the Sun from the Moon's Apogee; so is the  
 Minutes and Second (reduced into Seconds) taken out of  
 the following Table by help of the Sun's mean Anom-  
 ally, to the second Equation of the Moon in Second.

	Operation.	f. ° ' "
Sun's true Place	—	9 22 23 46
Moon's Apogee first time Equated, sub. 1	—	7 53 51
<hr/>		
Dist. Sun from Moon's Apogee,		8 14 29 55
Double,	—	4 28 59 50
Complement,	—	1 1 00 10
	Now say,	0 . . "
As Radius	—	90 0 0 10.000000
To S. double Dist. $\odot$ a $\searrow$ Apog.	31	0 10 9.711874
So $3' 54''$ out of the following Table,	234	2.369216
To the second Equation in Seconds,	120	2.081090

This Equation must always be added to the first Equated  
 Place of the Moon, while her Apogee passes from the  
 Square of the Sun to the Conjunction; but is subtracted  
 from thence in the Transit of the Apogee from a Conjun-  
 ction, to a Quadrature. That is, if the Distance of the Sun  
 from the Moon's Apogee be

Signs  $\left\{ \begin{array}{l} 0 1 2 6 7 8 \text{ subtr.} \\ 3 4 5 9 10 11 \text{ add.} \end{array} \right.$

So in the Example before us, the second Equation  $2'$  is  
 subtracted from  $2 5. 23^{\circ} 27' 23''$ , and there remains  $2 5.$   
 $23^{\circ} 25' 23''$  the Moon's Place equated the second time.

Note, always in your Work, reserve the Logar. Sine of  
 the double Distance of the Sun from the  $\searrow$  Apogee; for  
 you will have Occasion to use it in finding the  $\searrow$  present  
 Eccentricity in the 10th Precept following.



A Table of the proportional Part of the  
second Equation of the Moon; with the  
Logarithm.

○Ano.	Eq.	Sign 0 Logar.	Eq.	Sign 1 Logar.	Eq.	Sign 2 Logar.	○Ano.
	"		"		"		
0	214	2.330414	217		221		30
1	214		217		221		29
2	214		218	2.338456	221		28
3	214		218		221		27
4	214		218		221		26
5	214		218		222	2.346353	25
6	214		218		222		24
7	214		218		222		23
8	215	2.332438	218		222		22
9	215		218		222		21
10	215		219	2.340444	222		20
11	215		219		222		19
12	215		219		222		18
13	215		219		223	2.348305	17
14	215		219		223		16
15	215		219		223		15
16	216	2.334454	219		223		14
17	216		219		223		13
18	216		219		223		12
19	216		220	2.342423	223		11
20	216		220		223		10
21	216		220		224	2.350248	9
22	216		220		224		8
23	216		220		224		7
24	217	2.336460	220		224		6
25	217		220		224		5
26	217		220		224		4
27	217		221	2.344392	224		3
28	217		221		224		2
29	217		221		225	2.352182	1
30	217		221		225		0
Sign 11			Sign 10			Sign 9	



A Table of the Proportional Part of the  
second Equation of the Moon, with the  
Logarithm, continued.

☉Ano.	Eq.	Sign 3 Logar.	Eq.	Sign 4 Logar.	Eq.	Sign 5 Logar.	☽Ano.
	"		"		"		
0	225		228		232		30
1	225		229	2.359836	232		29
2	225		229		232		28
3	225		229		233	2.367356	27
4	225		229		233		26
5	225		229		233		25
6	226	2.354108	229		233		24
7	226		229		233		23
8	226		229		233		22
9	226		230	2.361728	233		21
10	226		230		233		20
11	226		230		234	2.369216	19
12	226		230		234		18
13	226		230		234		17
14	227	2.356026	230		234		16
15	227		230		234		15
16	227		230		234		14
17	227		231	2.363612	234		13
18	227		231		234		12
19	227		231		235	2.371068	11
20	227		231		235		10
21	227		231		235		9
22	228	2.357935	231		235		8
23	228		231		235		7
24	228		231		235		6
25	228		232	2.365488	235		5
26	228		232		235		4
27	228		232		236	2.372912	3
28	228		232		236		2
29	228		232		236		1
30	228		232		236		0

Sign 3

Sign 7

Sign 6



## 8. To find the Third Equation of the Moon.

This Equation depends upon the Distance of the Sun from the Moon's Nodes, and is greatest in the Octants, and is then  $47''$ , but in the Syzигias and Quadratures nothing: Therefore from the Sun's true Place, subtract the Place of the Node first Equated, and say, As Radius, to the Sine of the double Distance of the Sun from the next Syzигia, or Quadrature, so is  $47''$ , to the Equation required.

	<i>Operation.</i>	<i>s. ° ' "</i>	
Sun's true Place.	— —	9 22 23 46	
Node first Equated subtr.	— —	10 25 43 30	
		<hr/>	
Diff. Sun from the Node,		10 26 40 16	
Double,	— —	9 23 20 32	
Complement,	— —	2 6 39 28	
	Now say,	0 ' "	
As Radius	— —	90 0 0 10.000000	
To S. double Dist of ☉ à ♁		66 39 28	9.962917
So is the greatest Equation	— —	47	1.672098
To the present Equation	— —	43	1.635015

This Equation is added to the Moon's Place Equated the second Time, whilst the Nodes pass from the Sun's Conjunction to the Quadratures of the same; and is subtracted in the Transit from the Quadratures to Conjunction. And, according to the Theory, I have framed the following Table, which gives this Equation by Inspection.



## A Table of the third Equation of the Moon.

Enter this Table with the Distance of the Sun from the Node, and you have the Equation to be apply'd according to its Title.

⊙ ° 30	Signs $\frac{2}{3}$ Add	Signs $\frac{2}{3}$ Add	Signs $\frac{2}{3}$ Add	⊙ ° 15
	' "	' "	' "	
0	0	0	41	30
1	0	1	41	29
2	0	3	42	28
3	0	5	43	27
4	0	6	43	26
5	0	8	44	25
6	0	10	44	24
7	0	11	45	23
8	0	13	45	22
9	0	14	46	21
10	0	16	46	20
11	0	17	46	19
12	0	19	47	18
13	0	20	47	17
14	0	22	47	16
15	0	23	47	15
16	0	25	47	14
17	0	26	47	13
18	0	27	47	12
19	0	29	46	11
20	0	30	46	10
21	0	31	46	9
22	0	32	45	8
23	0	34	45	7
24	0	35	44	6
25	0	36	44	5
26	0	37	43	4
27	0	38	43	3
28	0	39	42	2
29	0	40	41	1
30	0	41	41	0
	Signs $\frac{5}{11}$ Sub.	Signs $\frac{4}{10}$ Sub.	Signs $\frac{3}{9}$ Sub.	



## 9. To find the Second Equation of the Moon's Apogee.

First, (as in the seventh Article hereof) if from the Sun's Place, you subtract the Place of the Apogee first Equated, the Remainder is called the Annual Argument, and in this Example is 8<sup>s</sup>. 14° 29' 55", which is demonstrated from the Theory thus:

Let T represent the Earth, TS a right Line joining the Earth and ☉; TACB, a right Line drawn from the Earth to the Place of the Moon's Apogee first Equated.

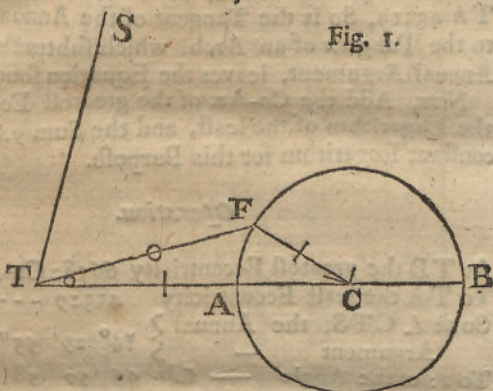


Fig. 1.

The Angle

STA, the Annual Argument, of the said Apogee = 74° 29' 55": TA the least Eccentricity 43319 of the Lunar Orbit; TB the greatest 66782. Bisect AB in C, and on the Center C, and Semidiameter AC = CB, Describe the Circle AFB; and make the Angle BCF equal to twice the Annual Argument 148° 59' 50", and draw TF, and CF, so shall TF be the Eccentricity at the time of the Question, and the Angle CTF the Equation of the Apogee sought. In which Triangle CTF are given TC, CF, and the Angle included, being the Complement of twice the Annual Argument to a Semicircle = 31° 0' 10" to find the Angle CTF, and the Side TF.

## Operation.

TB the greatest Eccentricity	66782	}	+ TA = TC
TA the least Eccentricity	43319		
Difference = AB	23463		
Half = AC = CF	11731½		Which



Which falls directly under the second Axiom of plain Trigonometry; for as the Sum of the Sides  $TC + CF$ , To their Difference  $= TA$ , So is the Tangent of half the Sum of the opposite Angles, to the Tangent of half their Difference, which subtracted from half the Sum of the Angles, gives the lesser Angle, *viz.*  $CTF$  the Equation sought. Which in short is, As the greatest Eccentricity  $TB$  66782, Is to the least Eccentricity  $TA$  43319, So is the Tangent of the Annual Argument, to the Tangent of an Arch, which subtracted from the Annual Argument, leaves the Equation sought.

*Note,* Add the Co-Ar. of the greatest Eccentricity to the Logarithm of the least, and the Sum 9.812019 is a constant Logarithm for this Purpose.

*Operation.*

As $TB$ the greatest Eccentricity	66782	Co-Ar.	5.175341	
To $TA$ the least Eccentricity	43319	----	4.636678	
So $\angle CTS$ , the Annual	}	74° 29' 55"	10.556970	
Argument				—
To $\angle$ of the Angle	—	sub.	66 50 58	10.368989

Rem:  $\angle CTF$ , the Equation 7 38 57

This Equation is to be added if the Annual Argument be 0, 1, 2, 6, 7 8 Signs, but subtracted if it be 3, 4, 5, 9, 10, 11, to or from the Place of the Apogee first Equated, the Sum or Difference is the Place of the Apogee a second time Equated: which in this Example is *1f.*  $15^{\circ} 32' 48''$ ; according to which Theory I have calculated the following Table.



A Table of the second Equation of the Moon's Apogee.

Ann. Argu.	Signs $\frac{2}{3}$ Add			Signs $\frac{1}{7}$ Add			Signs $\frac{2}{3}$ Add			Ann. Argu.
	o	'	"	o	'	"	o	'	"	
0	0	0	0	9	28	8	11	40	16	30
1	0	21	4	9	42	20	11	30	55	29
2	0	42	9	9	56	9	11	20	30	28
3	1	3	11	10	9	25	11	8	59	27
4	1	24	11	10	22	10	10	56	23	26
5	1	45	7	10	34	21	10	42	42	25
6	2	5	59	10	45	59	10	27	53	24
7	2	26	46	10	57	2	10	12	00	23
8	2	47	28	11	7	29	9	55	1	22
9	3	8	3	11	17	18	9	36	58	21
10	3	28	31	11	26	27	9	17	50	20
11	3	48	50	11	34	57	8	57	38	19
12	4	8	59	11	42	45	8	36	24	18
13	4	28	59	11	49	51	8	14	8	17
14	4	48	48	11	56	12	7	50	54	16
15	5	8	24	12	1	48	7	26	41	15
16	5	27	48	12	6	37	7	1	31	14
17	5	46	58	12	10	34	6	35	29	13
18	6	5	54	12	13	50	6	8	35	12
19	6	24	34	12	16	11	5	40	49	11
20	6	42	58	12	17	40	5	12	26	10
21	7	1	4	12	18	15	4	43	18	9
22	7	18	52	12	17	56	4	13	30	8
23	7	36	20	12	16	41	3	43	7	7
24	7	53	28	12	14	29	3	12	14	6
25	8	10	15	12	11	18	2	40	53	5
26	8	26	38	12	7	9	2	9	10	4
27	8	42	39	12	1	58	1	37	9	3
28	8	58	15	11	55	46	1	4	53	2
29	9	13	24	11	48	33	0	32	29	1
30	9	28	8	11	40	16	0	0	00	0
	Signs $\frac{5}{11}$ Sub.			Signs $\frac{4}{10}$ Sub.			Signs $\frac{2}{9}$ Sub.			



## 10. To find the present Eccentricity of the Moon.

Here are in the same Triangle TCF, given as before, with the Angle CTF just now found, to find the Side TF, the present Eccentricity. *Gregory Astron.* pag. 546.

Operation.

As <i>f.</i> $\angle$ CTF, the Equa.	} 7° 38' 57" Co-Ar.	0.875799
Apog. —	}	
To the Side CF (always the same)	} 11731 $\frac{1}{2}$ -----	4.069354
To <i>f.</i> $\angle$ TCF, Double Ano Ar —	} 31 00 10 -----	9.711874
To Side TF the Eccentricity —	} 45397 -----	4.657027

## 11. To find the Mean Anomaly of the Moon.

From the Moon's Place the third time Equated, subtract the Place of the Moon's Apogee the second time Equated, and the Remainder is the Mean Anomaly of the Moon at that Time.

Operation.

	<i>f.</i> ° ' "
From $\triangleright$ place the third time Equated	2 23 24 40
Subtract the Place Apogee second time Equ.	1 15 32 48
	<hr style="width: 100%;"/>
Remains $\triangleright$ Mean Anomaly	1 7 51 52

## 12. To find the Elliptic, or fourth Equation of the Moon.

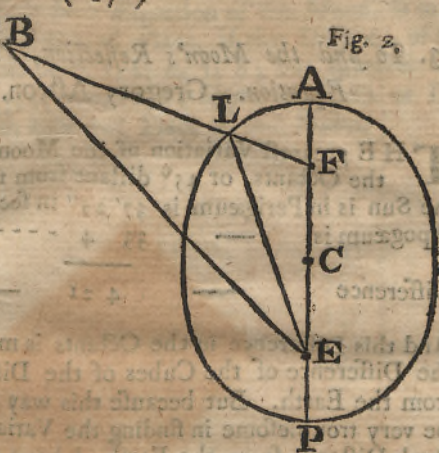
In the adjacent Figure make LB = LE, and join BE; then is FL + LE = AP. Make the Angle AFL equal to the Mean Anomaly 37° 51' 52 $\frac{1}{2}$ ; The Angle FEL the true Anomaly, and the Angle FLE the Elliptic Equation being the Double of the Angle FBE.

To



Fig. 2.

To AC the Mean  
 Distance of the  
 Moon from the  
 Earth = 1000000  
 add CE = CF  
 45397 the present  
 Eccentricity, the  
 Sum is 1045397 =  
 to AE the Apoge-  
 on Distance: again  
 from AC 1000000  
 take CE = CF  
 45397 the remain-  
 der is 954603 =  
 AF = EP the  
 Perigeon Distance.  
*Gregory's Astron.*  
 Vol. 1. pag. 389.



Now say,

As the Apogeeon Distance, to the Perigeon Distance: so  
 is the Tangent of half her Mean Anomaly, to the Tangent  
 of half the true Anomaly; which subtracted from half the  
 Mean Anomaly, and the Remainder doubled gives the  
 Prosthaphæresis or Elliptic Equation sought; which is to  
 be subtracted from the third Equated Place of the Moon,  
 if the Mean Anomaly be 0, 1, 2, 3, 4, 5 Signs; but added, if  
 it be 6, 7, 8, 9, 10, 11, the Sum or Difference is the fourth  
 Equated Place of the Moon.

Operation.

As AE	—	—	1045397	Co. Ar.	3.980719
To EP	—	—	954603	—	5.979822
			9		"
So <i>t.</i> of half Mean Anomaly	18	55	56	-	9.535301
To <i>t.</i> of half the true Anomaly	17	23	37	-	9.495842
Remains the Angle FBE	-	1	32	19	
Doubled is = ∠ FLE Equation	3	4	38	Subtract.	

D

13. To



13. To find the Moon's Reflection, Variation, or fifth Equation. Gregory Astron. pag. 548.

THE greatest Variation of the Moon is when she is in the Octants, or  $45^\circ$  distant from the Sun, and when the Sun is in Perigæum is  $37' 25''$  in seconds 2245"  
 Apogæum is ——— 33 4 ----- 1984  
 Difference ——— 4 21 ——— 261

And this Difference in the Octants is made reciprocally as the Difference of the Cubes of the Distances of the Sun from the Earth. But because this way of reasoning would be very troublesome in finding the Variation answering the Sun's Distance from the Earth, I have calculated the following Table, which enter with the Sun's Mean Anomaly, and you have the Variation answering. Then from the fourth Equated Place of the Moon, subtract the true Place of the Sun, which double, and say, As Radius to Sine of the double Distance of the Moon from the Sun: so is this Variation in respect of the Sun's Distance from the Earth, to the Variation of the  $\Delta$  in respect of the  $\odot$  at that Time.

## Operation.

4 Eq. pl. $\Delta$	—————	—————	2 20 20 2
$\odot$ place	—————	—————	9 22 23 46
Dist. $\Delta$ à $\odot$	—————	—————	4 27 56 16
Double	—————	—————	9 25 52 32
Complement	—————	—————	2 4 7 28

	Now say,	0 ' "
As Radius	—————	90 0 0 - 10.000000
To $f$ . Double	—————	64 7 28 - 9.954119
So $\odot$ Vari.	—————	-- 2225 - - 3.347330
To $\Delta$ Vari.	—————	2002 - - 3.301449

60) 2002 (33-22

This







A Table of the proportional Part of the 5<sup>th</sup> Equation, or Variation of the Moon.

Anom. $\odot$	Sign 0			Sign 1			Sign 2			Anom. $\odot$
	'	"	'''	'	"	'''	'	"	'''	
0	33	4	0	33	47	30	34	31	0	30
1	33	5	27	33	48	57	34	32	27	29
2	33	6	54	33	50	24	34	33	54	28
3	33	8	21	33	51	51	34	35	21	27
4	33	9	48	33	53	18	34	36	48	26
5	33	11	15	33	54	45	34	38	15	25
6	33	12	42	33	56	12	34	39	42	24
7	33	14	9	33	57	39	34	41	9	23
8	33	15	36	33	59	6	34	42	36	22
9	33	17	3	34	0	33	34	44	3	21
10	33	18	30	34	2	0	34	45	30	20
11	33	19	57	34	3	27	34	46	57	19
12	33	21	24	34	4	54	34	48	24	18
13	33	22	51	34	6	21	34	49	51	17
14	33	24	18	34	7	48	34	51	18	16
15	33	25	45	34	9	15	34	52	45	15
16	33	27	12	34	10	42	34	54	12	14
17	33	28	39	34	12	9	34	55	39	13
18	33	30	6	34	13	36	34	57	6	12
19	33	31	33	34	15	3	34	58	33	11
20	33	33	0	34	16	30	35	0	0	10
21	33	34	27	34	17	57	35	1	27	9
22	33	35	54	34	19	24	35	2	54	8
23	33	37	21	34	20	51	35	4	21	7
24	33	38	48	34	22	18	35	5	48	6
25	33	40	15	34	23	45	35	7	15	5
26	33	41	42	34	25	12	35	8	42	4
27	33	43	9	34	26	39	35	10	9	3
28	33	44	36	34	28	6	35	11	36	2
29	33	46	3	34	29	33	35	13	3	1
30	33	47	30	34	31	0	35	14	30	0
	Sign 11			Sign 10			Sign 9			



Table of the Proportional Part of the 5<sup>th</sup> Equation, or Variation of the Moon, continu'd.

Anom. ☉	Sign 3			Sign 4			Sign 5			Anom. ☉
	1	"	///	1	"	///	1	"	///	
0	35	14	30	35	58	0	36	41	30	30
1	35	15	57	35	59	27	36	42	57	29
2	35	17	24	36	0	54	36	44	24	28
3	35	18	51	36	2	21	36	45	51	27
4	35	20	18	36	3	48	36	47	18	26
5	35	21	45	36	5	15	36	48	45	25
6	35	23	12	36	6	42	36	50	12	24
7	35	24	39	36	8	9	36	51	39	23
8	35	26	6	36	9	36	36	53	6	22
9	35	27	35	36	11	3	36	54	33	21
10	35	29	0	36	12	30	36	56	0	20
11	35	30	27	36	13	57	36	57	27	19
12	35	31	54	36	15	24	36	58	54	18
13	35	33	21	36	16	51	37	0	21	17
14	35	34	48	36	18	18	37	1	48	16
15	35	36	15	36	19	45	37	3	15	15
16	35	37	42	36	21	12	37	4	42	14
17	35	39	9	36	22	39	37	6	9	13
18	35	40	36	36	24	6	37	7	36	12
19	35	42	3	36	25	33	37	9	3	11
20	35	43	30	36	27	0	37	10	30	10
21	35	44	57	36	28	27	37	11	57	9
22	35	46	24	36	29	54	37	13	24	8
23	35	47	51	36	31	21	37	14	51	7
24	35	49	18	36	32	48	37	16	18	6
25	35	50	45	36	34	15	37	17	45	5
26	35	52	12	36	35	42	37	19	12	4
27	35	53	39	36	37	9	37	20	39	3
28	35	55	6	36	38	36	37	22	6	2
29	35	56	33	36	40	3	37	23	33	1
30	35	58	0	36	41	30	37	25	0	0
	Sign 8			Sign 7			Sign 6			



## Add to the Variation of the Moon.

☾ hr ○	$f_{\frac{5}{8}}$		$f_{\frac{4}{7}}$		$f_{\frac{3}{2}}$		☾ hr ○
	/'	"	/'	"	/'	"	
0	0	0	2	54	2	54	30
1	0	6	3	0	2	48	29
2	0	12	3	6	2	42	28
3	0	17	3	11	2	37	27
4	0	23	3	17	2	31	26
5	0	29	3	23	2	25	25
6	0	35	3	29	2	19	24
7	0	41	3	35	2	13	23
8	0	46	3	40	2	8	22
9	0	52	3	46	2	2	21
10	0	58	3	52	1	56	20
11	1	4	3	58	1	50	19
12	1	10	4	4	1	44	18
13	1	15	4	9	1	39	17
14	1	21	4	15	1	33	16
15	1	27	4	21	1	27	15
16	1	33	4	27	1	21	14
17	1	39	4	33	1	15	13
18	1	44	4	39	1	10	12
19	1	50	3	58	1	4	11
20	1	56	3	52	0	58	10
21	2	2	3	46	0	52	9
22	2	8	3	40	0	46	8
23	2	13	3	35	0	41	7
24	2	19	3	29	0	35	6
25	2	25	3	23	0	29	5
26	2	31	3	17	0	23	4
27	2	37	3	11	0	17	3
28	2	42	3	6	0	12	2
29	2	48	3	0	0	6	1
30	2	54	2	54	0	0	0
	$f_{\frac{5}{17}}$		$f_{\frac{4}{10}}$		$f_{\frac{3}{6}}$		



A Table of the Variation or  
5<sup>th</sup> Equation.

n ○	$\frac{1}{2}$ Add.		$\frac{1}{7}$ Add.		$\frac{1}{8}$ Add.		n ○
	'	"	'	"	'	"	
0	0	0	28	38	28	38	30
1	1	12	29	11	28	2	29
2	2	21	29	43	27	25	28
3	3	29	30	12	26	45	27
4	4	37	30	39	26	4	26
5	5	46	31	4	25	20	25
6	6	53	31	27	24	34	24
7	7	59	31	47	23	47	23
8	9	7	32	5	22	58	22
9	10	13	32	20	22	7	22
10	11	19	32	34	21	15	21
11	12	24	32	45	20	22	19
12	13	27	32	53	19	26	18
13	14	30	32	58	18	29	17
14	15	32	33	3	17	32	16
15	16	32	33	4	16	32	15
16	17	32	33	3	15	32	14
17	18	29	32	58	14	30	13
18	19	26	32	53	13	27	12
19	20	22	32	45	12	24	11
20	21	15	32	34	11	19	10
21	22	7	32	20	10	13	9
22	22	58	32	5	9	7	8
23	23	47	31	47	7	59	7
24	24	34	31	27	6	53	6
25	25	20	31	4	5	46	5
26	26	4	30	39	4	37	4
27	26	45	30	12	3	29	3
28	27	25	29	43	2	12	2
29	28	2	29	11	1	12	1
30	28	38	28	38	0	0	0

|  $\frac{1}{2}$  Sub. |  $\frac{1}{7}$  Sub. |  $\frac{1}{8}$  Sub. |



## 14. To find the Sixth Equation of the Moon.

OUR Author fixes this Equation at a mean Quantity  $2' 10''$ . This being known; from the Moon's Apogee the second time Equated, subtract the Apogee of the Sun, and Note the Remainder.

Also from the Place of the Moon the fifth time Equated, subtract the Sun's true Place, and Note this Remainder, add these two Remainders together, and work as below.

*Operation.*

Place Moon's Apogee second time Equated	1	15	32	48
Place Sun's Apogee subtract	3	8	13	55
Remainder	10	7	18	53
Place of the $\Delta$ the fifth time Equated	2	19	46	40
Sun's Place subtract	9	22	23	46
Remainder	14	27	22	54
Remainder add	10	7	18	53
Sum of these two Remainders	3	4	41	47
Complement	2	25	18	13

Now say,

As Radius	90	0	0	10.000000
To Sine Z Remainders	85	18	13	9.998539
So is $2' 10''$ in Seconds	130	2	113943	
To the Sixth Equation in Seconds	129	2	112482	

This Sixth Equation is to be subtracted from the 5th Equated Place of the Moon, if the aforefaid Sum of the two Remainders (or its Excess above 12 Signs) be less than a Semicircle, or Six Signs; but if it is more, it must be added, the Sum or Difference is the Moon's Longitude the sixth time Equated.



## 15. To find the Seventh Equation of the Moon.

This Equation Sir *Isaac Newton* has expressed by this mean Quantity  $2' 20''$ , which, he says, is increased and diminished, according to the Situation of the Lunar Apogæum  $54''$ ; that is, if the Lunar Apogee be join'd with the Sun's Apogæum, the Equation is then  $2' 20'' + 54'' = 3' 14''$ , and also when it is in the Sun's Syzigi-  
as; but when it is in the Sun's Quadratures, the afore-  
said Equation is to be diminished  $50''$ , and he makes the  
least Quantity  $1' 26''$ , that is  $2' 20'' - 54''$ . But when the  
Moon's Apogee and Sun's are in Opposition, he cannot  
determine (for want of Observation) whether the said E-  
quation is to be increased or diminished.

Therefore, from the sixth Equated Place of the Moon, subtract the Sun's true Place, and the Remainder is the Distance of the Moon from the Sun. Then say,

As Radius, To the Sine of this Distance of the Moon from the Sun, So is  $2' 20''$  in Seconds  $140''$ , To the Seventh Equation of the Moon: which must be subtracted from the sixth Equated Place, if the Distance of the Moon from the Sun be less than 6 Signs, but added when more; the Sum or Difference is the Moon's true Place in her Orbit.

## Operation.

Moon's Longitude the sixth time Equated	f.	°	'	''
Sun's true Place subtract				
Remains the Distance of the $\text{D} \grave{\text{a}} \odot$				
Complement				

## Now say,

As Radius	°	'	''
To f. Dist. $\text{D} \grave{\text{a}} \odot$	90	0	0
So $2' 20'' =$ Seconds	32	39	15
To $\times 15$ the seventh Equation sub.	140	2.146	128
	75	1.878	173

E

16. 70



16. To find the Moon's Latitude, and the Reduction from her Orbit to the Ecliptic.

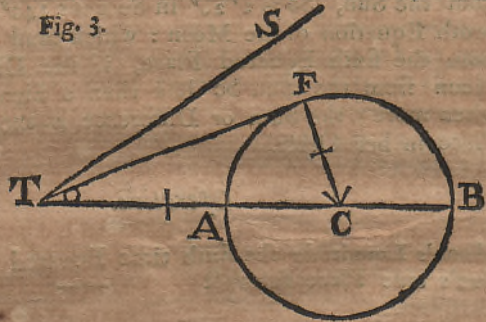
From the Sun's true Place, subtract the Place of the Moon's North Node first Equated, the Remainder is called the Annual Argument of the Node.

Operation.

	f. 0 1 0
Sun's true Place	9 22 23 46
North Node first Equated	10 25 43 30
	10 26 40 16
Rem. Annual Argument of the Node	10 26 40 16
Doubled is	9 23 20 32

From this Work, and from the adjacent Figure, the second Equation of the Node is thus constructed.

Let T represent the Earth, TS, a right Line joining the Earth and Sun: Let TA CB represent a Line drawn to the Place of the Moon's ascending Node Equated the first Time; and the Angle S T A the Annual Argument of the Node.



Let TA be taken in the same Ratio to AB as 56 to 3. Bisect BA in C, and on the Center C and Radius AC = CB draw the Circle AFB, and let the Angle BCF be equal to Double the Annual Argument in this Example  $133^{\circ} 20' 32''$  as found above; Draw TF, then will the Angle CTF be the second Equation of the Node Ascending. Now in the Oblique-Angled Plain Triangle TFC, there are known, the Sides TC, and CF, with the included



cluded Angle FCT, to find the Angle CTF the present Equation of the Node. For if TA be 56, then AB (by the Theory) is 3, and consequently  $AC = 1\frac{1}{2}$ . Therefore  $TC\ 56 + AC\ 1\frac{1}{2} = TC\ 57\frac{1}{2}$  then because all Lines drawn from the Center of a Circle to the Circumference thereof are equal, CF is also equal to  $AC = CB\ 1\frac{1}{2}$ , and the Angle TCF is the Complement of the Double of the Annual Argument  $66^{\circ} 39' 28''$  of the Node. This being the second Axiom of oblique-angled Plain Triangles. The first and second Terms in the Analogy are always the same, viz. the Sum of the Sides is ever 59, and their Difference 56: so that as the Sum of the Sides 59 is to their Difference 56, so is the Tangent of half the Sum of their opposite Angles (which is ever equal to the Annual Argument of the Node) To the Tangent of half their Difference; which subtracted from half the Sum of the two unknown Angles, leaves the lesser Angle CTF, which is the Equation of the Node.

*Operation.*

As the Sum of the Sides TC + CF Co-Ar.	59	—	8.229148
To their Difference	56	—	1.748188
	o	1	"
So t. of half Z opposite Angles	33	19	44—9.817961
To t. of half their Difference	31	58	15—9.795297
		—	
Remains Equation of the Node	1	21	29 Subtract.

This Equation is always to be added to the first Equated Place of the Node, if the Double of the Annual Argument be less than Six Signs; but subtracted, if more. Or with the Distance of the Sun from the Node enter the Table of its Equation (which I have made according to above Directions) and take out the Equation, which apply'd to the first Equated Place according to its Title gives the true Place of the Moon's Ascending Node. Note, add the Co-Ar. of the Logarithm of the two Sides, to the Logarithm of their Difference, and that Sum shall be a constant Logarithm 9.977336.



17. To find the Inclination of the Limit, Gregory  
Astron. pag. 550.

THE Inclination of the Moon's Orb with the Ecliptic, when the Nodes are in the Sun's Quadratures, is  $4^{\circ} 59' 35''$ . And when they are in the Syzигias  $5^{\circ} 17' 20''$ : therefore the greatest Inclination above the least is  $17' 45''$ ; and according to this Limitation I have made the Table of Inclination of the Limit, which enter with the Annual Argument of the Node, or Distance of the Sun from the Node, which in this Example is *10 f. 26° 40' 16''*, and take out the Inclination of the Limit, *12' 26''*.

18. From the true Place of the Moon in her Orbit, subtract the true Place of the Node last Equated, and the Remainder is the Argument of Latitude *3 f. 25° 21' 15''*.

19. With the Argument of Latitude take out of the Table the simple Latitude, as also the Increment or Parts to be added, and reserve them till anon. *Greg. 551.*

20. Before we can find the Moon's true Latitude, we must find the proportional Part of the Increment thus. As the greatest Inclination of the Moon's Orb and Ecliptic above the least, *viz. 17' 45''*.

To the present Increment as found by Precept 19.  
So is the present Inclination of the Limit as found by the 17, To the proportional Part of the Increment, which added to the Simple Latitude, as found by the 19, and it gives the Moon's true Latitude.

21. With the Argument of Latitude enter the Table of Reduction, and take out the Reduction, and also the Excess; then, for the proportional Part of the Excess by the Logistical Logarithms, say,

As the greatest Inclination of the Moon's Orb above the least, *viz. 17' 45''*,

Is to the present Excess:

So is the present Inclination of the Limit as found by the 17 above,

To



To the proportional Part of the Excess, which added to the simple Reduction, gives the true.

22. According to the Title of the Table of Reduction, apply the Reduction to the Moon's Orbit Place, and you will have the Moon's true Place in the Ecliptic: as you may the better perceive by tracing the following Examples.

The



## The first Example of the Sun and Moon's Place.

Equal Time	Long. ☉	Anom. ☉		
	f. ° ' "	f. ° ' "		
Anno 1729	9 20 56 32	6 12 42 37		
January 1	59 8	59 8		
Mean Mot.	9 21 55 40	6 13 41 45	Anom.	} ☉
Equat. add	28 6	9 21 55 40	Long.	
☉ true Plac.	9 22 23 46	3 8 13 55	Apoq.	

Equal Time	Long. ♃	Apoq. ♃	Node ♃	
	f. ° ' "	f. ° ' "	f. ° ' "	
Anno 1729	2 10 19 39	1 7 42 20	10 25 48 59	
January 1	13 10 35	6 41	3 11	
Mean Mot.	2 23 30 14	1 7 49 1	10 25 45 48	
1 Equation	— 2 51	+ 4 50	— 2 18	
☉ Equated 1	2 23 27 23	1 7 53 51	10 25 43 30	Node Equat.
2 Equation f.	— 2 0	+ 7 38 57	9 ☉ 22 23 46	Sun's Place.
☉ Equated 2	2 23 25 23	1 15 32 48	10 26 40 16	☉ à ☉.
3 Equation f.	— 43	9 ☉ 22 23 46	9 23 20 32	Double.
☉ Equated 3	2 23 24 40	8 14 29 55	2 6 39 28	Complement
4 Equation f.	— 3 4 38	4 28 59 50	— 1 21 29	2d Equat. sub.
☉ Equated 4	2 20 20 2	1 1 00 10	10 24 22 1	Nod. true Pl.
5 Equation f.	— 33 22	Eccentricity	— — — 45397	
☉ Equated 5	2 29 46 40	Mean Anom.	— — — 1 f. 7° 51' 52"	
6 Equation f.	— 2 9	Inclination of the Limit.	— — — 12 26	
☉ Equated 6	2 19 44 31	Simple Latitude	— — — 4 30 37	
7 Equation f.	— 1 15	Increment	— — — — 16 1	
☉ in her Orb.	2 19 43 16	Excefs	— — — — — 36	
North Nod. f.	10 24 22 1	As the Greatest	17' 45" Co-Ar. 4709	
Arg. Latitud.	3 25 21 15	To Increment	16 1 — — 5736	
Tr. Lat. N.D.	4 41 50	So Inclination	12 26 — — 6836	
Red. add	+ 5 29	To Increment	11 13 — — 7281	
Eclipt. Place	2 19 48 45			



## Example 2.

Equal Time	Long. ☉	Anom. ☉	
Anno 1729	f. ° 1 "	f. ° 1 "	
January 2	9 20 56 32	6 12 42 37	
Mean Mot.	1 58 17	1 58 17	
Equat. add	9 22 54 49	6 14 40 54	Anom. } ☉
☉ true Place	30 5	9 22 54 49	Long. } ☉
	9 23 24 54	3 8 13 55	Apog. } ☉

Equal Time	Long. ♃	Apog. ♃	Node ♃	
Anno 1729	f. ° 1 "	f. ° 1 "	f. ° 1 "	
January 2	2 10 19 39	1 7 42 20	10 25 48 59	
Mean Mot.	26 21 10	13 22	6 21	
1 Equation	3 6 40 49	1 7 55 42	10 25 42 38	
2 Equation f.	3 3	5 10	2 27	
3 Equation f.	3 6 37 46	1 8 0 52	10 25 40 11	Node Equat.
4 Equation f.	1 54	9 23 24 54	9 23 24 54	Sun's Place.
5 Equation f.	3 6 35 52	8 15 24 2	10 27 44 43	☉ à 38.
6 Equation f.	42	5 0 48 4	9 25 29 26	Double.
7 Equation f.	3 6 35 10	0 29 11 56	2 4 30 34	Complement.
8 Equation f.	3 55 38	7 16 43	1 20 1	ad Equat. sub.
9 Equation f.	3 2 39 32	1 15 17 35	10 24 20 10	Nod. true Pl.
10 Equation f.	24 34			
11 Equation f.	3 2 14 58	Eccentricity	— — —	45 173
12 Equation f.	2 5	Mean Anom.	— — —	1 21 17 35
13 Equation f.	3 2 12 53	Simple Lat.	— — —	3 56 21
14 Equation f.	51	Inclination	— — —	12 43
15 Equation f.	3 2 12 2	Increment	— — —	13 58
16 Equation f.	10 24 20 10	Excess	— — —	0 46
17 Equation f.	4 7 51 52	As the Greatest	17' 45" Co-AR.	4709
18 Equation f.	4 6 22	To Excess	— 0 46 — —	18935
19 Equation f.	Reduct. add	So Inclinat.	— 12 43 — —	6738
20 Equation f.	6 53	To Excess add	0 33 — —	20382
21 Equation f.	3 2 18 55			



## SECTION II.

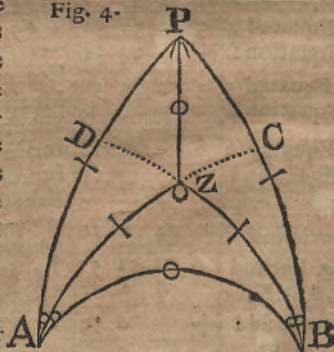
*A Problem of the Sphere.*

**G**IVEN, Two Altitudes of the Sun upon any Azimuth, with the Time between them, to find the Latitude of the Place.

*Example.* Suppose July 17, at 10 $\frac{1}{2}$  in the Forenoon, I observe the Sun's Altitude  $52^{\circ} 59'$ ; and at 2 in the Afternoon the same Day I observe it  $49^{\circ} 51'$ ; What is the Latitude of the Place of Observation?

In the following Figure, A represents the Time and Place of the first Observation, B the second, P the North Pole, Z the Zenith, in which Triangles are given, AP  $70^{\circ} 58'$  the Complement of the Sun's Declination, or its Distance from the North Pole, at the Time of the first Observation; BP  $71^{\circ} 1'$  the Sun's Distance from the North Pole at the Time of the second Observation, with the included Angle at P, *viz.* APB,  $52^{\circ} 30'$  the Time between the two Observations: Also in the Triangle AZB, there are given AZ  $37^{\circ} 1'$  the Complement of the Sun's Altitude, or its Distance from the Zenith at the Time of the first Observation, and BZ  $40^{\circ} 9'$  the Complement of the Sun's Altitude at the Time of the second Observation, to find ZP the Distance of the Zenith from the Pole equal to the Complement of the Latitude of the Place of

Fig. 4.



Observation. But before we can find ZP, we must find the Side AB, and the Angles at A and B. This may be effected two Ways; first, as I have shewed in Problem



blem 16 of my Doctrine of the Sphere. The other Way is wrought by the versed Sines, as I shall shew following. First in the Triangle A P B, to find the Side A B. This may be done by letting fall a Perpendicular from the Angle A upon B P, as A C; or else B D from the Angle B upon the Side A P, for both will produce the same Thing.

## Operation.

As Cr. P B	—	71	1	9.536561	}
To Radius	—	90	0	10.000000	
So C f. $\angle$ D P B	—	52	30	9.784447	
To t. D P -- Sub.	—	60	32	10.247886	
From A P	—	70	58		
Remains A D	—	10	26		
As C f. D P	—	60	32	Co. Ar. 0.308108	}
To C f. A D	—	10	26	9.992759	
To C f. B P	—	71	1	9.512275	
To C f. A B	—	49	26	9.813142	

Or by letting fall the Perpendicular A C.

As Cr. A P	—	70	58	9.537792	}
To Radius	—	90	0	10.000000	
So C f. $\angle$ A P C	—	52	30	9.784447	
To t. P C. Sub.	—	60	28	10.246655	
From P B	—	71	1		
Remains C B	—	10	33		
As C f. P C	—	60	28	Co. Ar. 0.307242	}
To C f. C B	—	10	33	9.992596	
So C f. A P	—	70	58	9.513375	
To C f. A B	—	49	26	9.813213	

Secondly, by the Tables of Versed Sines.

To solve this Case without the Help of the Perpendicular, has been long since hinted by *Gunter*, *Speidel*, *Gollibrand*, and *Collins*; (see *Collins's Sector on a Quadrant*, Pag. 88.) and since them by *Sir Jonas Moore*, *Harris*, &c.



*Rule.*

As the Cube of the Radius, Is to the Rectangle of the Sines of the Comprehending Sides:

So is the Square of the Sine of half the Angle contain'd, To half the Difference of the verfed Sines of the Third, and of the Ark of Difference between the two including Sides. Which is thus:

Double the Logarithm Sine of half the Angle given, and thereto add the Logarithm Sines of the contain'd Sides, and from the Left-hand of the Sum, dash out or reject 3, for the Cube of the Radius, so there rests the Logarithm of half the Difference of those two verfed Sines, which half Difference doubled and added to the verfed Sine of the Difference of the Logarithms of the containing Sides, gives the N. verfed Sine of the Side sought.

*Operation.*

Given Angle APB	52° 30'	{ half = 26° 15' }		
		{ Double Sine }	19.291412	
Given Sides	{	AP	70 58 - Sine	9.975583
		BP	71 1 - Sine	9.975713
Sum is f. of	10 4		9.242708	
Natural Sine of	10 4	is	1747939	
Doubled is			3495878	
Diff. of the Sides	0 3	Verfed Sine is add.	4	
Sum is N. Verfed f.	49 26	= Side AB	3495882	

Secondly, for the Angle ABP.

As f. AB	49 26	Co-Ar.	0.119387	} Or it may be found by the 9th Problem of my Doctrine of the Sphere, &c.
To f. ∠ APB	52 30		9.899467	
So f. AP	70 58		9.975583	
To f. ∠ ABP	80 51		9.994437	



## 3. For the Angle B A P.

As $f. AB$	$49^{\circ} 26'$	Co-Ar.	0.119387	}
To $f. \angle APB$	$52 30$	-----	9.899467	
So $f. PB$	$71 1$	-----	9.975713	
To $f. \angle BAP$	$80 57$	-----	9.994567	

## 4. For the Angle Z A B. By Prob. 9. of my Astron.

Side subtend. the required $\angle ZB$	$49^{\circ} 9'$	$AB$	$49^{\circ} 26'$
	$\frac{1}{2} 20$	$AZ$	$37 1$
	$\frac{1}{2} 6$		$12\frac{1}{2}$
		X	$12 25$
	$Z$	$26 17$	$\frac{1}{2} 6 12\frac{1}{2}$
	X	$13 52$	

$f. AZ$	$37^{\circ} 1'$	Co-Ar.	0.220369	}
$f. AB$	$49 26$	Co-Ar.	0.119387	
$f. Z$	$26 17$	-----	9.646218	
$f. X$	$13 52$	-----	9.379601	
Sum Logarithms	-----		19.365575	
Sine	$28 48$	-----	9.6827875	
Double is	$57 36$	$\angle ZAB$	sub.	
$BAP$	$80 57$	from		

$ZAP$   $23 21$  Rem.

## 5. For the Angle ZBA.

Side subtending the required	}	$37^{\circ} 1'$	$BZ$	$40^{\circ} 9'$
Angle is $AZ$		$18 30\frac{1}{2}$	$BA$	$49 26$
		$\frac{1}{2} 4$	$38\frac{1}{2}$	
			X	$9 17$
		$Z$	$23 9$	$\frac{1}{2} 4 38\frac{1}{2}$
		X	$13 52$	

$F 2$

$f. BZ$



<i>f</i> BZ	40	9	Co-Ar.	--	0.190581	} The three Angles at the Zenith are thus;
<i>f</i> BA	49	26	Co-Ar.	--	0.119387	
<i>f</i> Z	23	9			9.594547	
<i>f</i> X	13	52			9.379601	
Z of the			Logarithms	--	19.284116	
Sine of	26	1			9.642058	A Z P 143° 0'
						B Z P 132 56
						A Z B 84 4

Double 52 2 =  $\angle$  ZBA.

A B P 80 51

Z 360 00

Proves the work right.

Z B P 28 49

6. Lastly, for the Side ZP, being the Complement of the Latitude of the Place of Observation.

As *f*.  $\angle$  ZPB 30° 0 Co Ar. 0.301029

To *f*. — ZB 40 9 — — — 9.809419

So *f*.  $\angle$  ZBP 28 49 — — — 9.683055

To *f*. — ZP 38 26 — — — 9.793503

From — 90 0

Remains — 51 34, the Latitude of the Place North.

And thus may the Latitude of the Place be found without the Meridian Altitude, by taking the Sun's Altitude twice in the Forenoon, or twice in the Afternoon; or by the Moon alone, or by the Moon and any Star, or else by the Altitude of two fixed Stars or Planets, for their Difference of right Ascensions shall be the Angle at the Pole; and then the Things given and required in the Triangle, are the same as above, which needs no Example.

### S E C T. III.

*The Use of the Tables of the Satellites of Jupiter and Saturn.*

**T**HIS Part of Astronomy was entirely unknown to the Ancients, till about the Year 1610, when *Galileus* in *Italy* first discover'd that *Jupiter* was environ'd with four Moons or Satellites; and in *Germany* by *Simon Marius*, by Help of the Telescope, without which,



which, by Reason of *Jupiter's* Splendor, and their small Distance from him, they are not to be discerned. Thus, they being discovered, it put our Moderns upon examining their Motions, and framing a Theory, which is now sufficiently done; and I can now (when *Jupiter* is not too near the Sun) shew them upon Demand to any one that is minded to be satisfied of the Truth hereof. The Circum-saturnials are in Number Five, which by reason of their great Distance from the Sun, and the Smallness of their Bodies, are not to be seen but by the Help of very long Telescopes; the diligent *Cassini* was the first that saw the 1, 2, 3 and 5 Satellites, with a Telescope of 17 Feet, about the End of *October* in the Year 1671. But the 4th of *Saturn's* Moons was first discover'd by Mr. *Hugens*, Anno 1655. These Satellites of *Jupiter*, as well as of *Saturn*, are called *Secondary Planets, Moons, or Concomitants*; for they constantly keep close to their respective Primaries, and always attend upon them in their Circulations round the Sun; and in the mean time each of them performs his proper Revolution round his proper Primary: The Earth indeed has only the Moon to keep her company, who never forsakes her in her annual Course round the Sun; and while she attends upon us, she performs proper Circulations of her own round the Earth, in the Space of near a Month. The Satellites of *Saturn* have also been seen in *England* by means of that Telescope which was given to the Royal Society by the *Dutch* Astronomer, Mr. *Hugens*, whose Length is 125 Feet, by which nearly a perfect Theory of their Motions are settled; and from those Observations and Theory, I have formed these Tables of their Motions. To these Satellites belong peculiar Phenomena, which are these:

1. That they cannot be seen with the naked Eye.
2. That they cannot always be seen by them that look at them thro' a Telescope.
3. They always appear in a right Line with their Primary, some to the Right, and some to the Left-Hand; at other times, all to the Right, or all to the Left-Hand.
4. That they are continually changing their apparent Distances from their Primaries, seeming at one time to be near, and at another time further removed.
5. That as they are viewed by us from the Earth, they seem to go sometimes to the East, and at other times



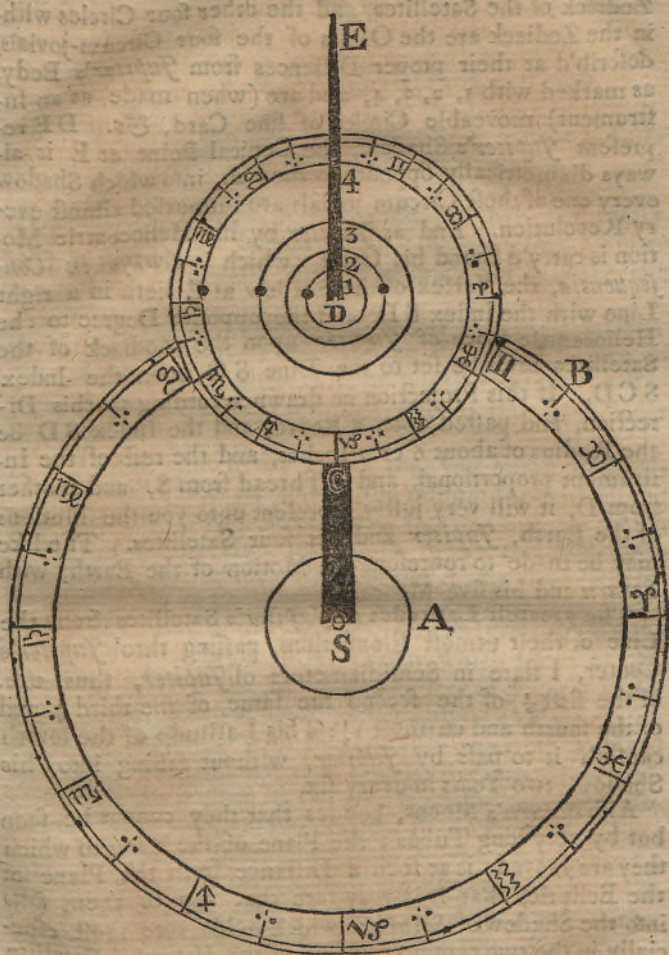
times to move to the West. 6. That when you look at them with two convex Glasses, they always appear on the contrary Side of their Primary to what they really are. 7. That Regard is to be had to the different Situation of the Earth, and the Place of the Observer, from these Phænomena of their Places; for such a Phænomenon is sooner seen when the Earth is nearer to it, than when it is removed from it. See the Theory of *Jupiter* in my System of the Planets, demonstrated Plate 10. Figure 4. 8. That when  $\text{h}$  is in the Nodes of the Satellites  $\text{M} \times 21^\circ$ , the Sun is in the Plane of the Ring, from those Places; therefore the right Ascension and Declination is computed. 9. As far as can be discovered by us at so great a Distance, the Orbit of these Satellites appearing little or nothing eccentric; and by comparing the Periods of their Motions with the Periods of our Moon round the Earth, and the Periods of *Jupiter* and *Saturn* with the Period of the Earth round the Sun; the Inequalities in the Motions of these Satellites, may be deriv'd from the Inequalities in the Moon's Motion describ'd above in her Theory. These being their chief Phænomena.

We well know that the Motion of these Satellites are disturbed by the Sun acting upon them, and also by each other's Attraction, and by the Attraction of their Superiours or Primary Planets; but these Things being more physical than astronomical, I shall not trouble my Reader with them at this Time, but refer them to Sir *Isaac Newton's* Philosophy, and *Gregory's* Elements.

The third Satellite from *Jupiter* is of the first Magnitude, the first of that next to his Body is the second Magnitude, the second Satellite is of the third, and the fourth or outmost is of the fourth Magnitude. By these Directions it will not be difficult for an Observer to distinguish one Satellite from another. And the *Hugean* or fourth of *Saturn's* Satellites, is the greatest of his Guards.

As the  $\text{D}$  (the Earth's Satellite) is carry'd round the Sun along with the Earth in a Year, so are also the Satellites of *Jupiter* and *Saturn* carry'd round the Sun in the times of their respective Revolutions of their Primaries. And for a Demonstration hereof, I have framed the following Diagram after the Nature of my Lunar Instrument, which will make their Motions very plain and easy; in which





let S represent the Sun, A the Earth's Orb, B *Jupiter's* Orb, divided into the 12 Signs of the Zodiack, the Index C is moveable, and lieth under the Orbs of the Satellites, and terminates at D, on which *Jupiter* at D moves round the Sun, and carries his four Satellites along with him; the outmost Circle from *Jupiter* at D represents the Zodiack marked with the 12 Signs thereof, which I call the  
Zo.



Zodiack of the Satellites, and the other four Circles within the Zodiack are the Orbits of the four Circum-jovials, describ'd at their proper Distances from *Jupiter's* Body, as marked with 1, 2, 3, 4, and are (when made as an Instrument) moveable Circles of fine Card, &c. DE represent *Jupiter's* Shadow, the Conical Point at E is always diametrically opposite to the Sun, into which Shadow every one of these Circum-jovials are immerf'd almost every Revolution. And as *Jupiter* by his Heliocentric Motion is carry'd round his Orbit, which is always in *Consequentia*, the Vertex of the Shadow at E lieth in a right Line with the Index SD, and the opposite Degree to the Heliocentric Place of *Jupiter* upon the Zodiack of the Satellites must be set to the Line SC upon the Index SCD. If this Projection be drawn according to this Direction, and pasted upon a Board, and the Index SD be the Radius of about 6 or 7 Inches, and the rest of the Instrument proportional, and a Thread from S, and another from D, it will very justly represent unto you the Motions of the Earth, *Jupiter* and his four Satellites. The like may be made to represent the Motion of the Earth, with *Saturn* and his five Moons.

The greatest Latitudes of *Jupiter's* Satellites from the Line of their utmost Elongations, passing thro' *Jupiter's* Center, I state in Semidiameters of *Jupiter*, thus, viz. of the first  $\frac{1}{2}$ , of the second the same, of the third  $\frac{1}{4}$ , and of the fourth and outmost  $1\frac{1}{2}$ : This Latitude of the fourth causeth it to pass by *Jupiter*, without falling into his Shadow, two Years in every six.

Also *Saturn's* Moons, besides that they cannot be seen but by very long Tubes; the Plane of the Ring in which they are moved, is at such a Distance from the Plane of the Ecliptic, that it is very rare that any of them falls into the Shadow of *Saturn*: which holds true most especially in the two remotest, whereof the *Hugenian* Satellite is one, whose Place I shall shew how to calculate by and by.



## P R E C E P T I.

*Shewing how to calculate the Distances of Jupiter's Satellites, and to distinguish one from another.*

**M**Y Design is here only to shew the ingenious Observer how to find at what Distance from *Jupiter* each Satellite appears, that so he may not mistake one from another.

1. From the Tables in my compleat System of Astronomy to the given Time, calculate the Heliocentric Place of *Jupiter*, as is there shew'd, from which subtract the Radical Place  $5^{\circ} 13' 22'' 57''$ , and reserve the Remainder till anon.

2. To the same given Time, that you found *Jupiter's* Heliocentric Place to, collect the middle Motions of the Satellite from its respective Table into one Sum, to which Sum, add always (after the Radix 1684) the Difference between the Heliocentric Place of the Radix, and the Heliocentric Place at the given Time just now reserv'd; this Sum is the Place of the Satellite at the given Time, to answer our present Purpose. From which Place of the Satellite, subtract the Heliocentric Place of *Jupiter* at the given Time, what remains is the Distance of the Satellite from *Jupiter*; with this Distance enter the Table of each Satellite's Distance from  $\mu$  under the respective Satellite, and there is given its Distance from the Center of *Jupiter* in Semidiameters and decimal Parts. Here note, That if the Distance of the Satellite from *Jupiter* be less than six Signs, it is then in Consequence of *Jupiter*; but if the Distance be more than six Signs, it is in Antecedence, as the said Table will more fully direct you.

*Example.*

Anno 1728, December 16<sup>d</sup> 9<sup>h</sup> 1' 48<sup>n</sup>, I observed with my  $13\frac{1}{2}$  Feet Tube, and two convex Glasses, the 2d, 3d, and 4th of *Jupiter's* Satellites (the first being at that time eclipsed) how were they by my Satellite Tables?



## Operation.

Equal Time	Long $\mu$				Anom. $\mu$			
	f.	o	'	"	f.	o	'	"
1728	1	26	56	17	7	16	48	53
Dec. 16 Biff.		29	10	48		29	9	37
H. 9			1	52			1	52
' "								
1 48								
M. Mot.	2	26	8	57	8	16	00	22
Equat. add.		5	25	56				
$\mu$ Helioc.	3	1	34	55				
Radix sub.	5	13	22	57				
$\mu$ à Radix	9	18	11	58				

Equal Time	Satellite 1.				Satellite 2.			Satellite 3.			Satellite 4.					
	f.	o	'	"	f.	'	"	f.	o	'	"	f.	o	'	"	
1728	11	2	39	1	2	8	15	54	6	27	46	57	3	24	19	14
Decemb.	8	17	32	1	11	21	22	34	7	8	19	29	11	26	57	38
Days 16 Biff.	7	7	53	54	9	11	57	23	4	13	59	12		5	17	40
Hours 9	2	16	16	38	1	7	59	3	18	50	16			8	3	29
Min. 1			8	29			4	13			2	6				54
Seconds 48			6	47			3	23			1	40				43
Sum —	5	14	36	50	0	19	42	30	7	8	59	40	4	4	39	38
$\mu$ à Radix	9	18	11	58	9	18	11	58	9	18	11	58	9	18	11	58
Satellite	3	02	48	48	10	7	54	28	4	27	11	38	1	22	51	36
$\mu$ Helioc.	3	1	34	55	3	1	34	55	3	1	34	55	3	1	34	55
Distance.	0	1	13	53	7	6	19	33	1	25	36	43	10	21	16	41

Hence, it appears by the Calculation, that the first Satellite was at the time of the Observation in the Shadow, because its Distance from *Jupiter's* Heliocentric Place was less than the Semiduration or Half-Stray in the Shadow  $9^{\circ}19'22''$ , and so of the other three.

The 2d and 4th were in Antecedence, and the third Satellite in Consequence of *Jupiter*: But appear'd on the contrary Side of him, for the Reasons above given.



2. The Satellites of *Saturn* may be found, as I have shewed in *Jupiter's*, with this Difference only, that as *Jupiter's* Satellites are immerfed into his Shadow, and *Saturn's* feldom are, that you work with *Saturn's* Geocentric Place, and you will have the Distance of his Moons from him, and also from each other; however, to fatisfy the Curious and more Inquifitive, I fhall here fhew the Method of *Hugens*, improv'd by Dr. *Halley*, in the *Philof. Transact.*

1. By the 8th Precept of my *Compleat System of Astronomy*, find the true Geocentric Place of *Saturn* to the time propofed; from which fubtract always the Place of the Apocronion (as they call it) 11 f.  $20^{\circ} 23' 48''$ , the Remainder is the Distance of *Saturn* from the Equinox of the Ring. To which Place or Remainder, find the right Afcenfion and Declination by the 2d and 3d Problems of my Book before cited; only here obferve, that you make Ufe of  $31^{\circ}$  for the Obliquity or Inclination of this Satellite, inftead of  $25^{\circ} 29'$ .

2. To the Right Afcenfion thus found, add the Place of the Apocronion 11 f.  $20^{\circ} 23' 48''$ , the Sum fhall be the Longitude of the Satellites Apogeon; then fay, As Radius, to the Sine of the Declination (found to the Obliquity  $31^{\circ}$ ) So is 8, To the greateft Latitude in Apogeon or Perigeon in the Parts of the Semidiameter of the Ring. Or fo is 18, to the Parts of the Semidiameter of *Saturn's* Globe.

3. To the given Time collect the middle Motion of the Satellite, and from it fubtract the Place of the Apocronion 11 20 23 48, the Remainder will be the mean Anomaly; with which, in the Table of the Moon's Equation in my *Compleat System of Astronomy*, take out the Equation answering thereto, and the Half thereof added or fubtracted to or from the mean Motion, according to the Title of the faid Lunar Table, gives the true Motion of the Satellite, from which fubtract the Apogeon as found in the laft Precept, and if the Remainder be more than fix Signs, the Satellite is Occidental; if lefs, Oriental: Then as Radius, to the Sine of the Remainder, fo is 8, to the Semidiameter of the Ring; or 18 to the Semidiameter of  $\eta$  Globe, that the Satellite is to the Eaftward or Weftward of the Center of *Saturn* accordingly.



4. As Radius, to Co-Sine of the said Remainder, so is the greatest Latitude from the Line of the Ansa, found by Precept 2. to the Latitude sought. Example, Anno 1657, May 19, New Stile, about 10 at Night, Mr. *Hugens* observed the 4 Satellites very near to  $\frac{1}{2}$  on the Western Side, and very little above the Line of the Ansa.

## Operation.

Anno 1657, May 9<sup>d</sup> 9<sup>h</sup> 40' P. M. at London.

		f.	o	'	"
Saturn's Geocentric Place	—	5	28	53	8
Apocronion sub.	—	11	20	23	48
<hr/>					
$\frac{1}{2}$ ab Equinox of the Ring	—	6	8	29	20
Right Ascension answering	—	6	7	17	0
Place Apogeeon	—	5	27	40	48
Declinat. South	—	4	21	0	0
Greatest Latitude in Apog.	$\frac{61}{100}$				
<hr/>					
Radix	—	1681	1	0	48
Mot. for Years sub.	—	24	8	29	35
<hr/>					
Remains the Radix	—	1657	4	1	13
May	—	6	9	14	6
<hr/>					
Days 9	—	6	23	11	33
Hours 9	—	8	27	59	
Min. 40	—	37	38		
<hr/>					
M. Motion	—	5	12	44	57
Apocr. sub.	—	11	20	23	48
<hr/>					
Anom.	—	5	22	21	9
<hr/>					
Equat. sub.	—	18	23		
<hr/>					
Long. Satell.	—	5	12	26	34
Apog. sub.	—	5	27	40	48
<hr/>					
Rem.	—	11	14	45	46
Comple.	—	0	15	14	14
True Latitude North	$\frac{12}{100}$				

Occident.

Before the Apogee.

Parts of the Ring.

Here



Here follows a Catalogue of the Observations from which I constructed these Tables of the Satellites of Jupiter and Saturn.

Anno 1657, May 19, N. S. about 10 at Night, Monf. Hugenſ, with his long Glaſſes, obſerved the 4 Satellites of Saturn a little above the Line of the Anſæ, and very near to Saturn on the Weſtern Side, that Planet was then in  $\mathbb{R}$   $28^{\circ} 53' 8''$ .

Anno 1658, Mar. 11, N. S. at 10 P. M. Monf. Hugenſ obſerved the ſame Satellite a little to the Eaſt of Saturn, and on the South-ſide of him.

Anno 1659, March 14, N. S. 12<sup>h</sup> P. M. at the Hague, Monf. Hugenſ obſerved the 4th Satellite about one Diameter of the Ring under Saturn: but it was gone ſo far to the Weſtward, that he concluded, that about four Hours before, or 7<sup>h</sup> 40' at London, it had been in Perigæo.

Anno 1659, March 22<sup>d</sup> 10<sup>h</sup> 45' P. M. this Satellite was a whole Diameter above the Line of the Anſæ, and the Perpendicular thereon fell nearly upon the Extremity of the Eaſtern Anſæ.

Anno 1682, November 13<sup>d</sup> 13<sup>h</sup> P. M. Dr. Halley obſerved Saturn's 4th Satellite in Perigæo on the North Side of him; and a Perpendicular let fall from it on the tranſverſe Diameter of the Ring, fell upon the Middle of the dark Spot of the following Anſæ.

Again, November 21<sup>d</sup> 16<sup>h</sup> 15' P. M. this Satellite was on the South Side, the Perpendicular on the Line of the Anſæ fell on the Middle of the dark Spot of the Weſtern Anſæ; and the ſame Night at 19<sup>h</sup>, the Perpendicular fell preciſely on the Center of Saturn; it was now in Apogæo.

Anno 1683, Jan. 24, at 8 at Night, he obſerved this Satellite in Apogæo, the Perpendicular on the Line of the Anſæ fell exactly on the Weſtern Limb of  $\mathbb{H}$ , and at 9<sup>h</sup> 30' the ſaid Perpendicular fell within the Globe more than half Way to the Center, and diſtant from the Line of the Anſæ towards the South, about one Diameter of the Ring.

Anno 1683, Feb. 9<sup>d</sup> 8<sup>h</sup> 10' P. M. this 4th Satellite was in Apogæo, and about one Diameter of Saturn's Ring to the South.

Anno



Anno 1714, April 4<sup>d</sup> 21<sup>h</sup> 30' N. S. Monf. *Cassini*, at the Royal Observatory in *France*, observed the innermost of  $\eta$  Satellites in its inferior Conjunction with *Saturn*, he was then in  $\mathbb{M}$  5<sup>o</sup> 23' 32" by my Tables, and the same time the second Satellite was in its superior Conjunction with *Saturn*.

Anno 1714, M. *Cassini* observed the third Satellite of *Saturn*, April 4<sup>d</sup> 10<sup>h</sup> N. S. to have newly pass'd its inferior Conjunction with *Saturn*, and a Perpendicular from it fell on the Extremity of the Western Ansa, so that at about 5<sup>h</sup> P. M. it was with the Center of *Saturn* then in  $\mathbb{M}$  5<sup>o</sup> 27' R, and consequently the Satellite was in  $\times$  5<sup>o</sup> 27'.

Anno 1715, March 25<sup>d</sup> 11<sup>h</sup> P. M. M. *Cassini* observed the Hugonian or 4th Satellite in Apogæo, and did immerge behind the Body of *Saturn*, he was then in  $\mathbb{M}$  20<sup>o</sup> 3' 33" R.

Anno 1714, May 6<sup>d</sup>, 12<sup>h</sup> P. M. *Cassini* observed the 5th or outmost Satellite in its superior Conjunction with  $\eta$  he being then Retrograde in  $\mathbb{M}$  4<sup>o</sup> 32' 45".

Anno 1718, April 21<sup>d</sup> 10<sup>h</sup> 4', Mr. *Pound* at *Wansted*, by Help of the Royal Society's 125 Feet Glass, observed the 3d and 4th Satellites of  $\eta$  in Apogæo, a little past their Conjunction with *Saturn*. The first was Northward of the Line of the Ansa; and therefore in the Apogæon Semicircle distant from the said Line  $\frac{3}{4}$  of *Saturn*'s Semidiameter, and about a Semidiameter of the Ring from the Western Ansa.

The 2d was a very little Southward of the Line of the Ansa (and therefore in the Perigeon Semicircle) above a Semidiameter of the Ring from the Western Ansa. The 3d, 1st and 2d, were in a straight Line.

And, April 22<sup>d</sup> 11<sup>h</sup> 5' P. M. the four innermost Satellites were all Eastward of  $\eta$ , the 2d and 4th in the Apogæon, and the 1st and 3d in the Perigeon Semicircle: The 5th or outmost Satellite was at this time near its greatest Elongation Eastward.

The Tables of the Satellites of *Jupiter* I have deduced from the Observations of Mr. *Flemsted*, confirmed with abundance of my own.



A Table of the Mean Motion of the Moon

Years	Long. D.				Apog. D.				Node D.			
	f.	o	'	"	f.	o	'	"	f.	o	'	"
1681	6	1	45	55	8	4	29	45	5	24	13	35
1682	10	11	8	58	9	15	9	35	5	4	53	52
1683	2	20	32	1	10	25	49	26	4	15	34	9
1684	6	29	55	4	0	6	29	18	3	26	14	26
1685	11	22	28	42	1	17	15	49	3	6	51	32
1686	4	1	51	45	2	27	55	39	2	17	31	49
1687	8	11	14	48	4	8	35	29	1	28	12	6
1688	0	20	37	51	5	19	15	20	1	8	52	23
1689	5	13	11	30	7	0	1	51	0	19	29	29
1690	9	22	34	34	8	10	41	42	0	0	9	46
1691	2	1	57	38	9	21	21	32	11	10	50	3
1692	6	11	20	42	11	2	1	23	10	21	30	20
1693	11	3	54	21	0	12	47	54	10	2	7	26
1694	3	13	17	24	1	23	27	45	9	12	47	43
1695	7	22	40	27	3	4	7	35	8	23	28	0
1696	0	2	3	31	4	14	47	26	8	4	8	17
1697	4	24	37	34	5	25	33	57	7	14	45	23
1698	9	4	0	12	7	6	13	48	6	25	25	40
1699	1	13	23	17	8	16	53	38	6	6	5	57
1700	5	22	46	21	9	27	33	28	5	16	46	14
1701	10	15	20	0	11	8	20	0	4	27	23	20
1702	2	24	43	3	0	18	59	50	4	8	3	37
1703	7	4	6	6	1	29	39	40	3	18	43	54
1704	11	13	29	9	3	10	19	31	2	29	24	11
1705	4	6	2	47	4	21	6	2	2	10	1	17
1706	8	15	25	50	6	1	45	52	1	20	41	34
1707	0	24	48	53	7	12	25	42	1	1	21	51



A Table of the Mean Motion of the Moon

Years	Long. D.				Apoq. D.				Node D.			
	f.	o	1	"	f.	o	1	"	f.	o	1	"
1708	5	4	11	56	8	23	5	34	0	12	2	8
1709	9	26	45	34	10	3	52	5	11	22	39	14
1710	2	6	8	38	11	14	31	55	11	3	19	31
1711	6	15	31	40	0	25	11	45	10	13	59	48
1712	10	24	54	45	2	5	51	36	9	24	40	5
1713	3	17	28	24	3	16	38	8	9	5	17	11
1714	7	26	51	27	4	27	17	58	8	15	57	28
1715	0	6	14	31	6	7	57	49	7	26	37	45
1716	4	15	37	36	7	18	37	40	7	7	18	2
1717	9	8	11	15	8	29	24	12	6	27	55	8
1718	1	17	34	19	10	10	4	2	5	28	35	25
1719	5	26	57	23	11	20	43	52	5	9	15	42
1720	10	6	20	27	1	1	23	43	4	19	55	59
1721	2	28	54	5	2	12	10	15	4	0	33	5
1722	7	8	17	8	3	22	50	5	3	11	13	22
1723	11	17	40	11	5	3	29	56	2	21	53	39
1724	3	27	3	14	6	14	9	46	2	2	33	56
1725	8	19	36	52	7	24	56	18	1	13	11	2
1726	0	28	59	55	9	5	36	8	0	23	51	19
1727	5	8	22	58	10	16	15	59	0	4	31	36
1728	9	17	46	1	11	26	55	49	11	15	11	53
1729	2	10	19	39	1	7	42	20	10	25	48	59
1730	6	19	42	42	2	18	21	10	10	6	29	16
1731	10	29	5	46	3	29	2	1	9	17	9	33
1732	3	8	28	50	5	9	41	50	8	27	49	50
1733	8	1	2	29	6	20	28	23	8	8	26	56
1734	0	10	25	33	8	1	8	13	7	19	7	13



in Years Current. 1735 to 1761 A

Years	Long. D.				Apog. D.				Node D.			
	f.	o	'	"	f.	o	'	"	f.	o	'	"
1735	4	19	48	37	9	11	48	4	6	29	47	30
1736	8	29	11	41	10	22	27	54	6	10	27	47
1737	1	21	45	20	0	3	14	26	5	21	4	53
1738	6	1	8	24	1	13	54	16	5	1	45	10
1739	10	10	21	28	2	24	34	7	4	12	25	27
1740	2	19	54	32	4	5	13	58	3	23	5	44
1741	7	12	28	10	5	16	0	30	3	3	42	50
1742	11	21	51	13	6	26	40	20	2	14	23	7
1743	4	1	14	17	8	7	20	11	1	25	3	24
1744	8	10	37	19	9	18	0	0	1	5	43	41
1745	1	3	10	58	10	28	46	32	0	16	20	47
1746	5	12	34	1	0	9	26	12	11	27	1	4
1747	9	21	57	5	1	20	6	13	11	7	41	21
1748	2	1	20	8	3	0	46	3	10	18	21	38
1749	6	23	53	47	4	11	32	34	9	28	58	44
1750	11	3	16	49	5	22	12	24	9	9	39	1
1751	3	12	39	54	7	2	52	15	8	20	19	18
1752	7	22	2	57	8	13	32	5	8	0	59	35
1753	0	14	36	36	9	24	18	38	7	11	36	41
1754	4	23	59	39	11	4	58	28	6	22	16	58
1755	9	3	22	42	0	15	38	19	6	2	57	15
1756	1	12	45	46	1	26	18	9	5	13	37	32
1757	6	5	19	26	3	7	4	41	4	24	14	38
1758	10	14	42	30	4	17	44	32	4	4	54	55
1759	2	24	5	34	5	28	24	22	3	15	35	12
1760	7	3	28	37	7	9	4	13	2	26	15	29
1761	11	26	2	15	8	19	50	45	2	6	52	35



A Table of the Mean Motion of the Moon

Years	Long. D.				Apog. D.				Node D.			
	f.	o	'	"	f.	o	'	"	f.	o	'	"
1762	4	5	25	18	10	0	30	35	1	17	32	52
1763	8	14	48	22	11	11	10	26	0	28	13	9
1764	0	24	11	25	0	21	50	16	0	8	53	26
1765	5	16	45	4	2	2	36	48	11	19	30	32
1766	9	26	8	8	3	13	16	38	11	0	10	49
1767	2	5	31	11	4	23	56	18	10	10	51	6
1768	6	14	54	15	6	4	36	19	9	21	31	23
1769	11	7	27	54	7	15	22	52	9	2	8	29
1770	3	16	50	57	8	26	2	42	8	12	48	46
1771	7	26	14	1	10	6	42	32	7	23	29	3
1772	0	5	37	4	11	17	22	23	7	4	9	20
1773	4	28	10	43	0	28	8	53	6	14	46	26
1774	9	7	33	46	2	8	48	43	5	25	26	43
1775	1	16	56	50	3	19	28	38	5	6	7	0
1776	5	26	19	53	5	0	8	26	4	16	47	17
1777	10	18	53	32	6	10	54	50	3	27	24	23
1778	2	28	16	35	7	21	34	47	3	8	4	40
1779	7	7	39	39	9	2	14	37	2	18	44	57
1780	11	17	2	42	10	12	54	28	1	29	25	14
1781	4	9	36	20	11	23	41	0	1	10	2	20
1782	8	18	59	23	1	4	20	51	0	20	42	37
1783	0	28	22	27	2	15	0	41	0	1	22	54
1784	5	7	45	30	3	25	40	32	11	12	3	11
1785	10	0	19	9	5	6	27	2	10	22	40	17
1786	2	9	42	12	6	17	6	53	10	3	20	34
1787	6	19	5	16	7	27	46	44	9	14	0	51
1788	10	28	28	19	9	8	26	35	8	24	41	8



in Years Current.

Years	Long. D.				Apog. D.				Node D.			
	f.	o	'	"	f.	o	'	"	f.	o	'	"
1789	3	21	1	57	10	19	13	7	8	5	18	14
1790	8	0	25	1	11	29	52	57	7	15	58	31
1791	0	9	48	4	1	10	32	48	6	26	38	48
1792	4	19	11	8	2	21	12	38	6	7	19	5
1793	9	11	44	47	4	1	59	10	5	17	56	11
1794	1	21	7	50	5	12	39	0	4	28	36	28
1795	6	0	30	54	6	23	18	50	4	9	16	45
1796	10	9	53	57	8	3	58	41	3	19	57	2
1797	3	2	27	36	9	14	45	12	3	0	34	8
1798	7	11	50	59	10	25	25	2	2	11	14	25
1799	11	21	13	43	0	6	4	53	1	21	54	42
1800	4	0	36	46	1	16	44	43	1	2	34	59
1801	8	23	10	25	2	27	31	15	0	13	12	5
1802	1	2	33	28	4	8	11	5	11	23	52	22
1803	5	11	56	32	5	18	50	56	11	4	32	39
1804	9	21	19	35	6	29	30	46	10	15	12	56
1805	2	13	53	14	8	10	17	18	9	25	50	2
1806	6	23	16	17	9	20	57	8	9	6	30	19
1807	11	2	39	21	11	1	36	59	8	17	10	36
1808	3	12	2	24	0	12	16	49	7	27	50	53
1809	8	4	36	3	1	23	3	20	7	8	27	59
1810	0	13	59	6	3	3	43	11	6	19	8	16
1811	4	23	22	10	4	14	23	2	5	29	48	33
1812	9	2	45	13	5	25	2	52	5	10	28	50
1813	1	25	18	51	7	5	49	24	4	21	5	56
1814	6	4	41	55	8	16	29	13	4	1	46	13
1815	10	14	4	58	9	27	9	4	3	12	26	30



A Table of the Mean Motion of the Moon, &amp;c.

Years	Long. D.				Apog. D.				Node D.			
	f.	o	'	"	f.	o	'	"	f.	o	'	"
1816	2	23	28	2	11	7	48	54	2	23	6	47
1817	7	16	1	41	0	18	35	26	2	3	43	53
1818	11	25	24	44	1	29	15	16	1	14	24	10
1819	4	4	47	48	3	9	55	6	0	25	4	27
1820	8	27	10	51	4	20	34	57	0	5	44	43
1821	1	6	44	30	6	1	21	30	11	16	21	50
1822	5	16	7	33	7	12	1	20	10	27	2	7
1823	9	25	30	37	8	22	41	10	10	7	42	24
1824	2	4	53	41	10	2	21	1	9	18	22	41
1825	6	27	27	19	11	14	7	32	8	28	59	47
1826	11	6	50	23	0	24	47	22	8	9	40	4
1827	3	16	13	26	2	5	27	12	7	20	20	21
1828	7	25	36	30	3	16	7	2	7	1	0	38
1829	0	18	10	9	4	26	53	34	6	11	37	44
1830	4	27	33	12	6	7	33	24	5	22	18	1
1831	9	6	56	15	7	18	13	15	5	2	58	18
1832	1	16	19	18	8	28	53	5	4	13	38	35
1833	6	8	52	57	10	9	39	37	3	24	15	41
1834	10	18	16	0	11	20	19	27	3	4	55	58
1835	2	27	39	4	1	0	59	18	2	15	36	15
1836	7	7	2	7	2	11	39	9	1	26	16	32
1837	11	29	35	46	3	22	25	41	1	6	53	38
1838	4	8	54	49	5	3	5	31	0	17	33	55
1839	8	18	21	53	6	13	45	21	11	28	14	12
1840	0	27	44	56	7	24	25	12	11	8	54	29
1841	5	20	18	34	9	5	11	45	10	19	31	36
1842	9	29	41	38	10	15	51	35	10	0	11	53



## The Mean Motion of the Moon.

JANUARY.												
Days	Long. D.				Apog. D.			Node D.				
	f.	o	'	"	o	'	"	o	'	"		
1	0	13	10	35	0	6	41	0	3	11		
2	0	26	21	10	0	13	22	0	6	21		
3	1	9	31	45	0	20	3	0	9	32		
4	1	22	42	20	0	26	44	0	12	43		
5	2	5	52	55	0	33	25	0	15	53		
6	2	19	3	30	0	40	6	0	19	4		
7	3	2	14	5	0	46	48	0	22	14		
8	3	15	24	40	0	53	29	0	25	25		
9	3	28	35	15	1	0	10	0	28	36		
10	4	11	45	50	1	6	51	0	31	46		
11	4	24	56	25	1	13	32	0	34	57		
12	5	8	7	0	1	20	13	0	38	8		
13	5	21	17	35	1	26	54	0	41	18		
14	6	4	28	10	1	33	35	0	44	29		
15	6	17	38	45	1	40	16	0	47	40		
16	7	0	49	20	1	46	57	0	50	50		
17	7	13	59	55	1	53	38	0	54	1		
18	7	27	10	30	2	0	19	0	57	11		
19	8	10	21	5	2	7	0	1	0	22		
20	8	23	31	40	2	13	41	1	3	33		
21	9	6	42	15	2	20	23	1	6	43		
22	9	19	52	50	2	27	4	1	9	54		
23	10	3	3	25	2	33	45	1	15	5		
24	10	16	14	0	2	40	26	1	16	15		
25	10	29	24	35	2	47	7	1	19	26		
26	11	12	35	10	2	53	48	1	22	37		
27	11	25	45	45	3	0	29	1	25	47		
28	0	8	56	20	3	7	10	1	28	58		
29	0	22	6	55	3	13	51	1	32	9		
30	1	5	17	31	3	20	32	1	35	19		
31	1	18	28	6	3	27	13	1	38	30		



## The Mean Motion of the Moon

F E B R U A R Y .											
Days	Long. D.				Agog. D.			Node D.			
	f.	o	'	"	o	'	"	o	'	"	
1	2	1	38	41	3	33	54	1	41	40	
2	2	14	49	16	3	40	35	1	44	50	
3	2	27	59	51	3	47	16	1	48	1	
4	3	11	10	26	3	53	57	1	51	12	
5	3	24	21	1	4	0	38	1	54	23	
6	4	7	31	36	4	7	19	1	57	33	
7	4	20	42	11	4	14	0	2	0	44	
8	5	3	52	46	4	20	41	2	3	54	
9	5	17	3	21	4	27	22	2	7	6	
10	6	0	13	56	4	34	4	2	10	16	
11	6	13	24	31	4	40	45	2	13	27	
12	6	26	35	6	4	47	26	2	16	37	
13	7	9	45	41	4	54	7	2	19	48	
14	7	22	56	15	5	0	48	2	22	59	
15	8	6	6	50	5	7	29	2	26	9	
16	8	19	17	26	5	10	10	2	29	20	
17	9	2	28	1	5	20	51	2	32	30	
18	9	15	38	36	5	27	32	2	35	41	
19	9	28	49	11	5	34	13	2	38	52	
20	10	11	59	46	5	40	54	2	42	2	
21	10	25	10	21	5	47	36	2	45	13	
22	11	8	20	56	5	54	17	2	48	23	
23	11	21	31	31	6	0	58	2	51	34	
24	0	4	42	6	6	7	39	2	54	45	
25	0	17	52	41	6	14	20	2	57	55	
26	1	1	3	16	6	21	1	3	1	6	
27	1	14	13	51	6	27	42	3	4	16	
28	1	27	24	26	6	34	23	3	7	27	
29	2	10	35	1	6	41	4	3	10	38	



( 55 )  
in Months and Days.

MARCH.

Com- mon	Long. D.				Apog. D.			Node D.			Bisex
	s.	o	l	"	o	l	"	o	l	"	
1	2	10	35	1	6	41	4	3	10	38	0
2	2	23	45	36	6	47	45	3	13	49	1
3	3	6	56	11	6	54	26	3	16	59	2
4	3	20	6	46	7	1	7	3	20	10	3
5	4	3	17	21	7	7	48	3	23	20	4
6	4	6	27	56	7	14	29	3	26	31	5
7	4	29	38	31	7	21	11	3	29	42	6
8	5	12	49	6	7	27	52	3	32	52	7
9	5	25	59	41	7	34	33	3	36	3	8
10	6	9	10	16	7	41	14	3	39	14	9
11	6	22	20	51	7	47	55	3	42	25	10
12	7	5	31	26	7	54	36	3	45	36	11
13	7	18	42	1	8	1	17	3	48	46	12
14	8	1	52	36	8	7	58	3	51	56	13
15	8	15	3	11	8	14	39	3	55	7	14
16	8	28	13	46	8	21	20	3	58	18	15
17	9	11	24	21	8	28	1	4	1	28	16
18	9	24	34	56	8	34	42	4	4	39	17
19	10	7	45	31	8	41	23	4	7	49	18
20	10	20	56	7	8	48	4	4	11	0	19
21	11	4	6	42	8	54	45	4	14	11	20
22	11	17	17	17	9	1	27	4	17	22	21
23	0	0	27	52	9	8	8	4	20	32	22
24	0	13	38	27	9	14	49	4	23	43	23
25	0	26	49	2	9	21	30	4	26	53	24
26	1	9	59	37	9	28	11	4	30	4	25
27	1	23	10	12	9	34	52	4	33	15	26
28	2	6	20	47	9	41	33	4	36	25	27
29	2	19	31	22	9	48	14	4	39	36	28
30	3	2	41	57	9	54	55	4	42	47	29
31	3	15	52	32	10	1	36	4	45	58	30



## The Mean Motion of the Moon

## A P R I L.

Com- mon	Long D.			Apog. D.			Node D.			Biflex	
	°	'	"	°	'	"	°	'	"		
1	3	29	3	7	10	8	17	4	49	8	0
2	4	12	13	42	10	14	58	4	52	11	1
3	4	25	24	17	10	21	39	4	55	29	2
4	5	8	34	52	10	28	20	4	58	40	3
5	5	21	45	27	10	35	2	5	1	51	4
6	6	4	56	2	10	41	43	5	5	1	5
7	6	18	6	38	10	48	24	5	8	12	6
8	7	1	17	12	10	55	5	5	11	22	7
9	7	14	27	47	11	1	46	5	14	33	8
10	7	27	38	22	11	8	27	5	17	44	9
11	8	10	48	57	11	15	8	5	20	54	10
12	8	23	59	32	11	21	49	5	24	5	11
13	9	7	10	7	11	28	30	5	27	16	12
14	9	20	20	42	11	35	11	5	30	26	13
15	10	3	31	17	11	41	52	5	33	37	14
16	10	16	41	52	11	48	33	5	36	48	15
17	10	29	52	27	11	55	14	5	39	58	16
18	11	13	3	2	12	1	55	5	43	9	17
19	11	26	13	37	12	8	36	5	46	15	18
20	0	9	24	12	12	15	18	5	49	31	19
21	0	22	34	47	12	21	59	5	52	41	20
22	1	5	45	22	12	28	40	5	55	52	21
23	1	18	55	47	12	35	21	5	59	2	22
24	2	2	6	32	12	42	2	6	2	13	23
25	2	15	17	7	12	48	43	6	5	54	24
26	2	28	27	42	12	55	24	6	8	34	25
27	3	11	38	17	13	2	5	6	11	45	26
28	3	24	48	52	13	8	46	6	14	56	27
29	4	7	59	27	13	15	27	6	18	6	28
30	4	21	10	3	13	22	8	6	21	17	29
31	5	4	20	37	13	28	49	6	24	27	30



( 57 )  
in Months and Days.

M A Y.

Com- mon	Long. D.				Apog. D.			Node D.			Bisex
	f.	o	l	"	o	l	"	o	l	"	
1	5	4	20	37	13	28	49	6	24	27	0
2	5	17	31	12	13	35	30	6	27	38	1
3	6	0	41	47	13	42	11	6	30	48	2
4	6	13	52	22	13	48	52	6	33	59	3
5	6	27	2	57	13	55	34	6	37	10	4
6	7	10	13	32	14	2	15	6	40	20	5
7	7	23	24	7	14	8	56	6	43	31	6
8	8	6	34	42	14	15	37	6	46	41	7
9	8	19	45	17	14	22	18	6	49	52	8
10	9	2	55	53	14	28	59	6	53	3	9
11	9	16	6	27	14	35	40	6	56	14	10
12	9	29	17	3	14	42	21	6	59	24	11
13	10	12	27	38	14	49	2	7	2	34	12
14	10	25	38	13	14	55	43	7	5	45	13
15	11	8	48	48	15	2	24	7	8	56	14
16	11	21	59	23	15	9	5	7	12	6	15
17	0	5	9	58	15	15	46	7	15	17	16
18	0	18	20	33	15	22	28	7	18	27	17
19	1	1	31	8	15	29	9	7	21	38	18
20	1	14	41	43	15	35	50	7	24	49	19
21	1	27	52	18	15	42	31	7	28	0	20
22	2	11	2	53	15	49	12	7	31	10	21
23	2	24	13	28	15	55	53	7	34	21	22
24	3	7	24	3	16	2	34	7	37	32	23
25	3	20	34	38	16	9	15	7	40	43	24
26	4	8	45	13	16	15	56	7	43	53	25
27	4	16	55	48	16	22	37	7	47	4	26
28	5	0	6	23	16	29	18	7	50	14	27
29	5	13	16	58	16	35	59	7	53	25	28
30	5	26	27	33	16	42	40	7	56	36	29
31	6	9	38	8	16	49	21	7	59	46	30



The Mean Motion of the Moon

J U N E.

Com- mon	Long D.				Apog. D.			Node D.			Bisex
	f.	o	1	"	o	1	"	o	1	"	
1	6	22	48	43	16	56	3	8	2	56	0
2	7	5	59	18	17	2	44	8	6	7	1
3	7	19	9	53	17	9	25	8	9	18	2
4	8	2	20	28	17	16	6	8	12	29	3
5	8	15	31	3	17	22	47	8	15	39	4
6	8	28	41	38	17	29	28	8	18	50	5
7	9	11	52	13	17	36	9	8	22	0	6
8	9	25	2	48	17	42	50	8	25	11	7
9	10	8	13	23	17	49	31	8	28	22	8
10	10	21	23	58	17	56	12	8	31	32	9
11	11	4	34	33	18	2	53	8	34	43	10
12	11	17	45	8	18	9	34	8	37	54	11
13	0	0	55	43	18	16	15	8	41	5	12
14	0	14	6	18	18	22	56	8	44	16	13
15	0	27	16	53	18	29	37	8	47	26	14
16	1	10	27	28	18	36	19	8	50	37	15
17	1	23	38	3	18	43	0	8	53	47	16
18	2	6	48	38	18	49	41	8	56	58	17
19	2	19	59	13	18	52	22	9	0	9	18
20	3	3	9	48	19	3	3	9	3	19	19
21	3	16	20	23	19	9	44	9	6	30	20
22	3	29	30	58	19	16	25	9	9	40	21
23	4	12	41	33	19	23	6	9	12	51	22
24	4	25	52	8	19	29	47	9	16	2	23
25	5	9	2	43	19	36	28	9	19	12	24
26	5	22	13	18	19	43	9	9	22	23	25
27	6	5	23	53	19	49	50	9	25	34	26
28	6	18	34	28	19	56	31	9	28	45	27
29	7	1	45	4	20	3	12	9	31	55	28
30	7	14	55	39	20	9	54	9	35	6	29
31	7	28	6	14	20	16	35	9	38	16	30



( 59 )  
in Months and Days.

JULY.

Com- mon	Long. D.				Apog. D.			Node D.			Bifex
	f.	o	/	"	o	/	"	o	/	"	
1	7	28	6	14	20	16	35	9	38	16	0
2	8	11	16	49	20	23	16	9	41	27	1
3	8	24	27	24	20	29	57	9	44	37	2
4	9	7	37	59	20	36	38	9	47	48	3
5	9	20	48	34	20	43	19	9	50	59	4
6	10	3	59	9	20	50	0	9	54	9	5
7	10	17	9	44	20	56	41	9	57	20	6
8	11	0	20	19	21	3	22	10	0	30	7
9	11	13	30	54	21	10	3	10	3	41	8
10	11	26	41	29	21	16	44	10	6	51	9
11	0	9	52	4	21	23	25	10	10	2	10
12	0	23	2	39	21	30	6	10	13	13	11
13	1	6	13	14	21	36	47	10	16	24	12
14	1	19	23	49	21	43	28	10	19	35	13
15	2	2	34	24	21	50	9	10	22	45	14
16	2	15	44	59	21	56	52	10	25	56	15
17	2	28	55	34	22	3	32	10	29	6	16
18	3	12	6	9	22	10	13	10	32	17	17
19	3	25	16	44	22	16	54	10	35	28	18
20	4	8	27	19	22	23	35	10	38	39	19
21	4	21	37	54	22	30	15	10	41	49	20
22	5	4	48	29	22	36	56	10	45	0	21
23	5	17	59	4	22	43	37	10	48	11	22
24	6	1	9	39	22	50	19	10	51	21	23
25	6	14	20	14	22	57	0	10	54	32	24
26	6	27	30	49	23	3	41	10	57	42	25
27	7	10	41	20	23	10	22	11	0	53	26
28	7	23	51	59	23	17	3	11	4	3	27
29	8	7	2	34	23	23	44	11	7	14	28
30	8	20	13	9	23	30	25	11	10	25	29
31	9	3	23	44	23	37	6	11	13	36	30



## The Mean Motion of the Moon

AUGUST.												
Com- mon	Long. D.				Agog. D.			Node D.			Eclix	
	s.	o	1	"	o	1	"	o	1	"		
1	9	16	34	19	23	43	47	11	16	47	0	
2	9	29	44	54	23	50	28	11	19	58	1	
3	10	12	55	29	23	57	9	11	23	8	2	
4	10	26	6	4	24	3	51	11	26	19	3	
5	11	9	16	39	24	10	32	11	29	29	4	
6	11	22	27	14	24	17	30	11	32	40	5	
7	0	5	37	49	24	23	54	11	35	51	6	
8	0	18	48	24	24	30	35	11	39	2	7	
9	1	1	58	59	24	37	16	11	42	12	8	
10	1	15	9	34	24	43	57	11	45	23	9	
11	1	28	20	9	24	50	38	11	48	35	10	
12	2	11	30	44	24	57	19	11	51	44	11	
13	2	24	41	19	25	4	0	11	54	54	12	
14	3	7	51	54	25	10	42	11	58	5	13	
15	3	21	2	29	25	17	23	12	1	15	14	
16	4	4	13	4	25	24	4	12	4	26	15	
17	4	17	23	40	25	30	45	12	7	36	16	
18	5	0	34	15	25	37	26	12	10	47	17	
19	5	13	44	50	25	44	7	12	13	58	18	
20	5	26	55	25	25	50	48	12	17	8	19	
21	6	10	6	0	25	57	29	12	20	19	20	
22	6	23	16	35	26	4	10	12	23	29	21	
23	7	6	27	10	26	10	51	12	26	40	22	
24	7	19	37	45	26	17	32	12	29	51	23	
25	8	2	48	20	26	24	13	12	33	1	24	
26	8	15	58	55	26	30	55	12	36	12	25	
27	8	29	9	30	26	37	36	12	39	23	26	
28	9	12	20	5	26	44	17	12	42	34	27	
29	9	25	30	40	26	50	58	12	45	44	28	
30	10	8	41	15	26	57	39	12	48	55	29	
31	10	21	51	50	27	4	20	12	52	5	30	



( 61 )  
in Months and Days.

SEPTEMBER.

Com- mon	Long. D.				Apog. D.				Node D.			Diffex
	f.	o	1	"	f.	o	1	"	o	1	"	
1	11	5	2	25	o	27	11	1	12	55	16	o
2	11	18	13	o	o	27	17	42	12	58	27	1
3	o	1	23	35	o	27	24	23	13	1	37	2
4	o	14	34	10	o	27	31	4	13	4	48	3
5	o	27	44	45	o	27	37	45	13	7	58	4
6	1	10	55	20	o	27	44	26	13	11	9	5
7	1	24	5	55	o	27	51	7	13	14	20	6
8	2	7	16	30	o	27	57	48	13	17	31	7
9	2	20	27	5	o	28	4	29	13	20	41	8
10	3	3	37	40	o	28	11	11	13	23	52	9
11	3	16	48	15	o	28	17	52	13	27	3	10
12	3	29	58	50	o	28	24	33	13	30	14	11
13	4	13	9	25	o	28	31	14	13	33	24	12
14	4	26	20	o	o	28	37	55	13	36	35	13
15	5	9	30	35	o	28	44	36	13	39	45	14
16	5	22	41	10	o	28	51	17	13	42	56	15
17	6	5	51	45	o	28	57	58	13	46	7	16
18	6	19	2	20	o	29	4	39	13	49	17	17
19	7	2	12	55	o	29	11	20	13	52	28	18
20	7	15	23	30	o	29	18	1	13	55	38	19
21	7	28	34	5	o	29	24	42	13	58	49	20
22	8	11	44	40	o	29	31	23	14	2	o	21
23	8	28	55	15	o	29	38	4	14	5	10	22
24	9	8	5	50	o	29	44	45	14	8	21	23
25	9	21	16	25	o	29	51	27	14	11	31	24
26	10	4	27	o	o	29	58	8	14	14	42	25
27	10	17	37	35	1	o	4	49	14	17	53	26
28	11	o	48	10	1	o	11	30	14	21	3	27
29	11	13	58	45	1	o	18	11	14	24	14	28
30	11	27	9	20	1	o	24	52	14	27	24	29
31	o	10	19	55	1	o	31	33	14	30	35	30



## The Mean Motion of the Moon

OCTOBER.

Com- mon	Long. D.				Apog. D.				Node D.			Biflex
	f.	o	l	"	f.	o	l	"	o	l	"	
1	0	10	19	55	1	0	31	33	14	30	25	0
2	0	23	30	30	1	0	38	14	14	33	46	1
3	1	6	41	5	1	0	44	55	14	36	56	2
4	1	19	51	40	1	0	51	36	14	40	7	3
5	2	3	2	15	1	0	58	17	14	43	17	4
6	2	16	12	50	1	1	4	58	14	46	28	5
7	2	29	23	26	1	1	11	39	14	49	39	6
8	3	12	34	1	1	1	18	20	14	52	50	7
9	3	25	44	36	1	1	25	2	14	56	0	8
10	4	8	55	11	1	1	31	43	14	59	11	9
11	4	22	5	46	1	1	38	24	15	2	21	10
12	5	5	16	21	1	1	45	5	15	5	32	11
13	5	8	26	56	1	1	51	46	15	8	43	12
14	6	1	37	31	1	1	58	27	15	11	53	13
15	6	14	48	6	1	2	5	8	15	15	4	14
16	6	27	58	41	1	2	11	49	15	18	15	15
17	7	11	9	16	1	2	18	30	15	21	26	16
18	7	24	19	51	1	2	25	11	15	24	36	17
19	8	7	30	26	1	2	31	52	15	27	47	18
20	8	20	41	1	1	2	38	33	15	30	57	19
21	9	3	51	36	1	2	45	14	15	34	8	20
22	9	17	2	11	1	2	51	55	15	37	19	21
23	10	0	12	46	1	2	58	36	15	40	29	22
24	10	13	23	21	1	3	5	17	15	43	40	23
25	10	26	33	56	1	3	11	58	15	46	50	24
26	11	9	44	31	1	3	18	39	15	50	1	25
27	11	22	55	6	1	3	25	21	15	53	12	26
28	0	6	5	41	1	3	32	2	15	56	22	27
29	0	19	16	16	1	3	38	43	15	59	33	28
30	1	2	26	51	1	3	45	24	16	2	43	29
31	1	15	37	26	1	3	52	5	16	5	54	30



( 63 )  
in Months and Days.

NOVEMBER.

Com- mon	Long. D.				Apog. D.				Node D.				Differ
	f.	o	'	"	f.	o	'	"	o	'	"		
1	1	28	48	1	1	3	58	46	16	9	5	0	
2	2	11	58	36	1	4	5	27	16	12	16	1	
3	2	25	9	11	1	4	12	8	16	15	26	2	
4	3	8	19	46	1	4	18	49	16	18	37	3	
5	3	21	30	21	1	4	25	30	16	21	48	4	
6	4	4	40	56	1	4	32	11	16	24	59	5	
7	4	17	51	31	1	4	38	53	16	28	9	6	
8	5	1	2	6	1	4	45	34	16	31	20	7	
9	5	14	12	41	1	4	52	15	16	34	31	8	
10	5	27	23	16	1	4	58	56	16	37	41	9	
11	6	10	33	51	1	5	5	37	16	40	52	10	
12	6	23	44	26	1	5	12	10	16	44	2	11	
13	7	6	55	1	1	5	18	59	16	47	13	12	
14	7	20	5	36	1	5	25	40	16	50	23	13	
15	8	3	16	11	1	5	32	21	16	53	34	14	
16	8	16	26	46	1	5	39	2	16	56	45	15	
17	8	29	37	21	1	5	45	43	16	59	55	16	
18	9	12	47	56	1	5	52	24	17	3	6	17	
19	9	25	58	31	1	5	59	5	17	16	16	18	
20	10	9	9	6	1	6	5	46	17	9	27	19	
21	10	22	19	41	1	6	12	27	17	12	38	20	
22	11	5	30	16	1	6	19	9	17	15	49	21	
23	11	18	40	51	1	6	25	50	17	18	59	22	
24	0	1	51	26	1	6	32	31	17	22	10	23	
25	0	15	2	2	1	6	39	12	17	25	21	24	
26	0	28	12	37	1	6	45	53	17	28	32	25	
27	1	11	23	12	1	6	52	34	17	31	42	26	
28	1	24	33	47	1	6	59	15	17	34	53	27	
29	2	7	44	22	1	7	5	56	17	38	3	28	
30	2	20	54	57	1	7	12	37	17	41	14	29	
31	3	4	5	32	1	7	19	18	17	44	25	30	



The Mean Motion of the Moon, &c.

DECEMBER.

Com- mon	Long. D.				Apog. D.				Node D.			Bifex
	f.	o	'	"	f.	o	'	"	o	'	"	
1	3	4	5	32	1	7	19	18	17	44	25	0
2	3	17	16	7	1	7	25	59	17	47	35	1
3	4	0	26	42	1	7	32	40	17	50	46	2
4	4	13	37	17	1	7	39	21	17	53	56	3
5	4	26	47	52	1	7	46	2	17	57	7	4
6	5	9	58	27	1	7	52	43	18	0	18	5
7	5	23	9	2	1	7	59	24	18	3	28	6
8	6	6	19	37	1	8	6	6	18	6	39	7
9	6	19	30	12	1	8	12	47	18	9	49	8
10	7	2	40	47	1	8	19	28	18	13	0	9
11	7	15	51	22	1	8	26	9	18	16	11	10
12	7	29	1	57	1	8	32	50	18	19	21	11
13	8	12	12	32	1	8	39	31	18	22	32	12
14	8	25	23	7	1	8	46	12	18	25	42	13
15	9	8	33	42	1	8	52	53	18	28	53	14
16	9	21	44	17	1	8	59	34	18	32	4	15
17	10	4	54	52	1	9	6	15	18	35	15	16
18	10	18	5	27	1	9	12	56	18	38	25	17
19	11	1	16	2	1	9	19	38	18	41	36	18
20	11	14	26	37	1	9	26	19	18	44	47	19
21	11	27	37	12	1	9	33	0	18	47	58	20
22	0	10	47	47	1	9	39	41	18	51	9	21
23	0	23	58	22	1	9	46	22	18	54	19	22
24	1	7	8	57	1	9	53	3	18	57	30	23
25	1	20	19	32	1	9	59	44	19	0	41	24
26	2	3	30	7	1	10	6	25	19	3	51	25
27	2	16	40	42	1	10	13	6	19	7	2	26
28	2	29	51	17	1	10	19	47	19	10	12	27
29	3	13	1	52	1	10	26	28	19	13	23	28
30	3	26	12	27	1	10	33	9	19	16	33	29
31	4	9	23	2	1	10	39	50	19	19	43	30



The Moon's mean Motion in Hours,  
Minutes and Seconds.

H	Long. D.			Apog. D.			Node D.	
	°	'	"	°	'	"	°	'
1	0	32	56	0	17		0	8
2	1	5	53	0	33		0	16
3	1	38	49	0	50		0	24
4	2	11	46	1	7		0	32
5	2	44	42	1	24		0	40
6	3	17	39	1	40		0	48
7	3	50	35	1	57		0	56
8	4	23	32	2	14		1	4
9	4	56	28	2	30		1	12
10	5	29	25	2	47		1	19
11	6	2	21	3	4		1	27
12	6	35	18	3	21		1	35
13	7	8	14	3	37		1	43
14	7	41	10	3	54		1	51
15	8	14	7	4	11		1	59
16	8	47	3	4	27		2	7
17	9	20	0	4	44		2	15
18	9	52	56	5	1		2	23
19	10	25	53	5	18		2	31
20	10	58	49	5	34		2	39
21	11	31	46	5	51		2	47
22	12	4	42	6	8		2	55
23	12	37	39	6	24		3	3
24	13	10	35	6	41		3	11
25	13	43	32	6	58		3	19
26	14	16	28	7	15		3	27
27	14	49	24	7	31		3	34
28	15	22	21	7	48		3	42
29	15	55	17	8	5		3	50
30	16	28	14	8	21		3	58



The Moon's mean Motion in Hours,  
Minutes and Seconds, continu'd.

H	Long. D.			Apog. D.		Node D.	
	°	'	"	'	"	'	"
31	17	28	14	8	38	4	6
32	17	34	7	8	54	4	14
33	18	7	3	9	11	4	22
34	18	39	59	9	28	4	30
35	19	12	55	9	45	4	38
36	19	45	52	10	2	4	46
37	20	18	48	10	19	4	54
38	20	51	45	10	36	5	2
39	21	24	41	10	52	5	10
40	21	27	38	11	8	5	18
41	22	30	34	11	25	5	26
42	23	3	31	11	42	5	34
43	23	36	27	11	59	5	42
44	24	9	24	12	16	5	50
45	24	42	20	12	32	5	58
46	25	15	17	12	48	6	6
47	25	48	13	13	5	6	14
48	26	21	10	13	22	6	22
49	26	54	6	13	39	6	30
50	27	27	3	13	56	6	38
51	27	59	59	14	13	6	46
52	28	32	56	14	30	6	54
53	29	5	52	14	46	7	1
54	29	38	49	15	2	7	8
55	30	11	45	15	19	7	16
56	30	44	42	15	36	7	24
57	31	17	38	15	53	7	32
58	31	50	34	16	10	7	40
59	32	23	31	16	26	7	48
60	32	56	27	16	43	7	56







A Table of the Simple Latitude of the Moon fitted to the least Inclination of its Orbit, with the Increments to the greatest Inclination.

Arg. L.	Latit.		Incre- ment Add.	Latit.		Incre- ment Add.	Latitude		Incre- ment Add.	Arg. L.
	0 N. A. 6 S. A.	1 "		1 N. A. 7 S. A.	2 "		2 N. A. 8 S. A.	3 "		
0	0 0 0	0 0	2 29 39	8 45	4 19 22	15 21	30			
1	0 5 14	0 11	2 34 9	9 1	4 21 58	15 30	29			
2	0 10 27	0 22	2 38 37	9 17	4 24 26	15 39	28			
3	0 15 40	0 41	2 43 1	9 33	4 26 52	15 47	27			
4	0 20 52	1 0	2 47 22	9 58	4 29 12	15 56	26			
5	0 26 4	1 19	2 51 41	10 4	4 31 27	16 4	25			
6	0 31 17	1 38	2 55 56	10 19	4 33 37	16 12	24			
7	0 36 28	1 56	3 00 8	10 34	4 35 43	16 19	23			
8	0 41 38	2 15	3 4 18	10 49	4 37 43	16 27	22			
9	0 46 48	2 34	3 8 23	11 4	4 39 38	16 34	21			
10	0 51 57	2 53	3 12 25	11 19	4 41 28	16 40	20			
11	0 57 5	3 11	3 16 24	11 33	4 43 13	16 46	19			
12	1 2 13	3 30	3 20 19	11 47	4 44 53	16 52	18			
13	1 7 18	3 48	3 24 11	12 1	4 46 27	16 57	17			
14	1 12 23	4 6	3 27 58	12 15	4 47 57	17 3	16			
15	1 17 27	4 24	3 31 42	12 29	4 49 21	17 8	15			
16	1 22 29	4 42	3 35 22	12 41	4 50 39	17 13	14			
17	1 27 29	5 0	3 38 58	12 54	4 51 53	17 18	13			
18	1 32 28	5 18	3 42 30	13 7	4 53 1	17 22	12			
19	1 37 26	5 36	3 45 58	13 20	4 54 4	17 25	11			
20	1 42 21	5 54	3 49 22	13 32	4 55 1	17 29	10			
21	1 47 14	6 12	3 52 42	13 44	4 55 53	17 32	9			
22	1 52 6	6 30	3 55 58	13 56	4 56 39	17 35	8			
23	1 56 56	6 57	3 59 9	14 8	4 57 20	17 37	7			
24	2 1 43	7 4	4 2 16	14 19	4 57 56	17 39	6			
25	2 6 28	7 21	4 5 18	14 30	4 58 26	17 41	5			
26	2 11 11	7 38	4 8 16	14 41	4 58 51	17 43	4			
27	2 15 52	7 54	4 11 9	14 51	4 59 10	17 44	3			
28	2 20 30	8 11	4 13 58	15 2	4 59 24	17 44	2			
29	2 25 7	8 28	4 16 42	15 11	4 59 32	17 45	1			
30	2 29 39	8 45	4 19 22	15 21	4 59 35	17 45	0			
	5 N. } 11 S. }	Defeen.	4 N. } 10 S. }	Defeen.	3 N. } 9 S. }	Defeen.				



A Table of Reduction, with the Excess above the least Inclination  $4^{\circ} 59' 35''$ .

Arg. L.	Signs <sup>0</sup> <sub>6</sub>			Exc.	Signs <sup>1</sup> <sub>7</sub>			Exc.	Signs <sup>2</sup> <sub>8</sub>			Exc.	Arg. L.
	Reduct. sub.	"	"		Reduct. sub.	"	"		Reduct. sub.	"	"		
0	0	0	0	0	5	40	42	5	40	42	30		
1	0	14	2	5	47	43	5	33	41	29			
2	0	27	4	5	53	44	5	25	40	28			
3	0	41	6	5	59	45	5	17	39	27			
4	0	55	8	6	4	46	5	10	38	26			
5	1	8	9	6	8	46	5	1	37	25			
6	1	22	11	6	13	46	4	52	36	24			
7	1	35	12	6	17	47	4	42	35	23			
8	1	48	13	6	20	47	4	33	34	22			
9	2	1	15	6	23	47	4	23	33	21			
10	2	14	17	6	26	47	4	13	31	20			
11	2	27	19	6	28	47	4	2	30	19			
12	2	40	20	6	30	48	3	51	29	18			
13	2	52	22	6	31	48	3	40	28	17			
14	3	4	23	6	32	48	3	29	26	16			
15	3	16	24	6	32	48	3	17	24	15			
16	3	28	26	6	32	48	3	5	23	14			
17	3	40	28	6	31	48	2	53	22	13			
18	3	51	29	6	30	48	2	40	20	12			
19	4	2	30	6	29	47	2	28	19	11			
20	4	12	31	6	26	47	2	15	17	10			
21	4	23	33	6	23	47	2	2	15	9			
22	4	33	34	6	20	47	1	49	13	8			
23	4	44	35	6	17	47	1	35	12	7			
24	4	52	36	6	12	46	1	22	11	6			
25	5	1	37	6	8	46	1	8	9	5			
26	5	9	38	6	3	46	0	55	8	4			
27	5	18	39	5	58	45	0	41	6	3			
28	5	26	40	5	52	44	0	27	4	2			
29	5	33	41	5	46	43	0	14	2	1			
30	5	40	42	5	40	42	0	0	0	0			
	Signs <sup>5</sup> <sub>11</sub>	Add.		Signs <sup>4</sup> <sub>10</sub>	Add.		Signs <sup>3</sup> <sub>9</sub>	Add.					



A Table of the Hourly Motions, Semi-diameters, and Horizontal Parallaxes of the Sun and Moon; the Sun's Horizontal Parallax being always 10".

Mean Anom of ☉ & ☾.		Hourly Motion ☉.		Hourly Motion ☾.		Mean Anom. of ☉ & ☾.	
f.	o.	1	11	1	11	f.	o.
0	0	2	23	29	37	12	0
0	6	2	23	29	38	11	24
0	12	2	23	29	41	11	18
0	18	2	23	29	47	11	12
0	24	2	23	29	55	11	6
1	0	2	24	30	7	11	0
1	6	2	24	30	21	10	24
1	12	2	24	30	35	10	18
1	18	2	24	30	51	10	12
1	24	2	25	31	10	10	6
2	0	2	25	31	31	10	0
2	6	2	26	31	54	9	24
2	12	2	26	32	18	9	18
2	18	2	27	32	43	9	12
2	24	2	27	33	7	9	6
3	0	2	28	33	33	9	0
3	6	2	28	33	59	8	24
3	12	2	29	34	26	8	18
3	18	2	29	34	54	8	12
3	24	2	30	35	21	8	6
4	0	2	30	35	46	8	0
4	6	2	31	36	9	7	24
4	12	2	31	36	30	7	18
4	18	2	32	36	50	7	12
4	24	2	32	37	9	7	6
5	0	2	32	37	27	7	0
5	6	2	33	37	42	6	24
5	12	2	33	37	55	6	18
5	18	2	33	38	4	6	12
5	24	2	33	38	9	6	6
6	0	2	33	38	10	6	0



A Table of the Hourly Motions, Semidiameters, and Horizontal Parallaxes of the Sun and Moon; the Sun's Horizontal Parallax being always  $10''$ : continued.

Mean Anom. of ☉ & ☾.		Apparen. Semedia. ☉.		Apparen. Semidia. ☾.		Horizont. Parallax. ☾.		Mean Anom. of ☉ & ☾.	
f.	°	'	"	'	"	'	"	f.	°
0	0	15	51	14	55	54	29	12	0
0	6	15	51	14	55	54	30	11	24
0	12	15	51	14	57	54	34	11	18
0	18	15	51	14	58	54	40	11	12
0	24	15	52	15	0	54	47	11	6
1	0	15	53	15	3	54	56	11	0
1	6	15	54	15	5	55	6	10	24
1	12	15	55	15	8	55	17	10	18
1	18	15	56	15	12	55	29	10	12
1	24	15	57	15	15	55	42	10	6
2	0	15	59	15	19	55	56	10	0
2	6	16	0	15	23	56	12	9	24
2	12	16	2	15	28	56	29	9	18
2	18	16	3	15	33	56	48	9	12
2	24	16	5	15	39	57	8	9	6
3	0	16	7	15	45	57	30	9	0
3	6	16	8	15	51	57	52	8	24
3	12	16	10	15	56	58	12	8	18
3	18	16	12	16	2	58	31	8	12
3	24	16	14	16	6	58	49	8	6
4	0	16	16	16	11	59	6	8	0
4	6	16	18	16	15	59	21	7	24
4	12	16	19	16	19	59	35	7	18
4	18	16	20	16	23	59	48	7	12
4	24	16	21	16	26	60	0	7	6
5	0	16	21	16	29	60	11	7	0
5	6	16	22	16	32	60	21	6	24
5	12	16	23	16	34	60	30	6	18
5	18	16	23	16	36	60	38	6	12
5	24	16	24	16	38	60	45	6	6
6	0	16	24	16	40	60	51	6	0



72 A Table of the Motion of the first Satellite of *Jupiter*.

Years Cur- rent-	Motion.				Years Com.	Motion.				Days.	Motion.			
	f.	o	'	"		f.	o	'	"		f.	o	'	"
1681	10	8	50	5	1	2	23	6	50	1	6	23	24	21
1701	3	18	8	28	2	5	16	13	40	2	1	16	48	42
1721	8	27	26	50	3	8	9	20	30	3	8	16	13	2
1722	11	20	33	40	4	5	25	51	40	4	3	3	37	23
1723	2	13	40	30	5	8	18	38	30	5	9	27	1	44
1724	5	6	47	20	6	11	12	5	20	6	4	20	26	5
1725	2	23	18	31	7	2	5	12	10	7	11	13	50	26
1726	5	16	25	21	8	11	21	43	21	8	6	7	14	47
1727	8	9	32	11	9	2	14	50	11	9	1	0	39	7
1728	11	2	39	1	10	5	7	57	1	10	7	24	3	28
1729	8	19	10	12	11	8	1	3	51	11	2	17	27	49
1730	11	2	17	2	12	5	17	35	1	12	9	10	52	10
1731	1	25	23	52	13	8	10	41	51	13	4	4	16	30
1732	4	18	30	41	14	11	3	48	41	14	10	27	40	52
1733	2	5	1	52	15	1	26	55	31	15	5	21	5	12
1734	4	28	8	42	16	11	13	26	42	16	0	14	29	33
1735	7	21	15	32	17	2	6	33	32	17	7	7	53	54
1736	10	14	22	22	18	4	29	40	22	18	2	1	18	15
1737	8	0	53	33	19	7	22	47	12	19	8	24	42	36
1738	10	24	0	23	20	5	9	18	23	20	3	18	6	57
1739	1	17	7	13	40	10	18	36	45	21	10	11	31	17
1740	4	10	14	3	60	3	27	55	8	22	5	4	55	38
1741	2	6	45	13	80	9	7	13	30	23	11	28	19	59
1742	11	23	16	24	100	2	16	31	53	24	6	21	44	20
1743	2	16	23	14	200	5	3	3	46	25	1	15	8	41
1744	5	9	30	4	300	7	19	35	39	26	8	8	33	2
1745	2	26	1	15	400	10	6	7	32	27	3	1	57	22
Jan.	0	0	0	0	500	0	22	39	25	28	9	25	21	43
Feb.	6	5	34	46	600	3	9	11	18	29	4	18	46	4
Mar.	4	0	56	30	700	5	25	43	12	30	11	12	10	25
Apr.	10	6	31	16	800	8	12	15	5	31	6	5	34	46
May	9	18	41	41	900	10	28	46	58					
June	3	24	16	27	1000	1	15	18	51					
July	3	6	26	52	2000	3	0	47	42					
Aug.	9	12	1	39	3000	4	15	56	33					
Sept.	3	17	36	25	4000	6	1	15	24					
Oct.	2	29	46	50	5000	7	16	34	15					
Nov.	9	5	21	36	6000	9	1	53	6					
Dec.	8	17	32	1	7000	10	17	11	57					



A Table of the Motion of the first Satellite of *Jupiter*. 73

H.	Motion.				/	o	/	"	'	o	/	"								
	f.	o	'	"									///	"	///	///	///	"	///	///
1	0	8	28	31	1	0	8	29	31	4	22	45								
2	0	16	57	2	2	0	16	57	32	4	31	13								
3	0	25	25	33	3	0	25	26	33	4	39	41								
4	1	3	54	3	4	0	33	54	34	4	48	10								
5	1	12	22	34	5	0	42	23	35	4	56	38								
6	1	20	51	5	6	0	50	51	36	5	5	7								
7	1	29	19	36	7	0	59	20	37	5	13	35								
8	2	7	48	7	8	1	7	48	38	5	22	4								
9	2	16	16	38	9	1	16	17	39	5	30	32								
10	2	24	45	9	10	1	24	45	40	5	39	1								
11	3	3	13	40	11	1	33	14	41	5	47	29								
12	3	11	42	10	12	1	41	42	42	5	55	58								
13	3	20	10	41	13	1	50	11	43	6	4	26								
14	3	28	39	12	14	1	58	39	44	6	12	55								
15	4	7	7	43	15	2	7	8	45	6	21	23								
16	4	15	36	14	16	2	15	36	46	6	29	52								
17	4	24	4	45	17	2	24	5	47	6	38	20								
18	5	2	33	16	18	2	32	33	48	6	46	49								
19	5	11	1	46	19	2	41	2	49	6	55	17								
20	5	19	30	17	20	2	49	30	50	7	3	46								
21	5	27	58	43	21	2	57	59	51	7	12	14								
22	6	6	27	19	22	3	6	27	52	7	20	43								
23	6	14	55	50	23	3	14	56	53	7	29	11								
24	6	23	24	21	24	3	23	24	54	7	37	40								
					25	3	31	53	55	7	46	8								
					26	3	40	21	56	7	54	37								
					27	3	48	50	57	8	3	5								
					28	3	57	18	58	8	11	34								
					29	4	5	47	59	8	20	2								
					30	4	14	15	60	8	28	31								

In Leap-Year, after *February*, add the Motion of a Day more to the rest. The Radix of this Satellite I have taken *December* 31 at Noon 1683. *Jupiter's* Heliocentric Place was then 5 f. 13° 22' 57". And the Half-Stay in the Shadow is 1<sup>h</sup> 6'. Mot. 9° 19' 22".



74 A Table of the Motion of the second Satellite of *Jupiter*.

Years Current	Motion.	Years Com.	Motion.	Days	Motion.
	f. o ' "		f. o ' "		f. o ' "
1681	3 7 35 50	1	8 11 24 51	1	3 11 17 30
1701	7 22 20 10	2	4 22 49 41	2	6 22 34 59
1721	0 7 4 30	3	1 4 14 31	3	10 3 52 29
1722	8 18 29 20	4	0 26 56 51	4	1 15 9 58
1723	4 29 54 11	5	9 8 21 41	5	4 26 27 28
1724	1 11 19 2	6	5 19 46 32	6	18 7 44 57
1725	1 4 1 22	7	2 1 11 22	7	11 19 2 27
1726	9 15 26 12	8	1 23 53 43	8	3 0 19 56
1727	5 26 51 3	9	10 5 18 33	9	6 11 37 26
1728	2 8 15 54	10	6 16 43 23	10	9 22 54 56
1729	2 0 58 14	11	2 28 8 14	11	1 4 12 25
1730	10 12 23 4	12	2 20 50 34	12	4 15 29 55
1731	6 23 47 55	13	11 2 15 25	13	7 26 47 24
1732	3 5 12 46	14	7 13 40 46	14	11 8 4 54
1733	2 27 55 6	15	3 25 5 6	15	2 19 22 23
1734	11 9 19 56	16	3 17 47 27	16	6 0 39 53
1735	7 20 44 47	17	11 29 12 17	17	9 11 57 23
1736	4 2 9 37	18	8 10 37 8	18	0 23 14 52
1737	3 24 51 58	19	4 22 1 59	19	4 4 32 22
1738	0 6 16 48	20	4 14 44 20	20	7 15 49 51
1739	8 17 41 39	40	8 29 28 40	21	10 27 7 21
1740	4 29 6 29	60	1 14 13 0	22	2 8 24 50
1741	4 21 48 50	80	5 28 57 20	23	5 19 42 20
1742	1 3 13 40	100	10 13 41 39	24	9 0 59 50
1743	9 14 38 31	200	8 27 23 19	25	0 12 17 19
1744	5 26 3 21	300	7 11 4 58	26	3 23 34 49
1745	5 18 45 42	400	5 24 46 38	27	7 4 52 18
Jan.	0 0 0 0	500	4 8 28 17	28	10 16 9 48
Feb.	8 20 2 16	600	2 22 9 56	29	1 27 27 17
Mar.	7 6 12 4	700	1 5 31 36	30	5 8 44 47
Apr.	3 26 14 20	800	11 19 33 15	31	8 20 2 16
May	9 4 29 7	900	10 3 14 45		
June	5 25 1 23	1000	8 16 56 24		
July	11 3 46 10	2000	5 3 52 48		
Aug.	7 23 48 26	3000	1 20 49 12		
Sept.	4 13 50 42	4000	10 7 45 36		
Oct.	9 22 35 29	5000	6 24 42 0		
Nov.	6 12 37 45	6000	3 11 38 24		
Dec.	11 21 22 31	7000	11 28 34 48		



A Table of the Motion of the Second Satellite of Jupiter. 75

H	Motion.				/	o	/	"	"	/	o	/	"
	f.	o	'	"									
1	o	4	13	14	1	o	4	13	31	2	10	50	
2	o	8	26	27	2	o	8	26	32	2	15	3	
3	o	12	39	41	3	o	12	40	33	2	19	17	
4	o	16	52	55	4	o	16	53	34	2	23	30	
5	o	21	6	9	5	o	21	6	35	2	27	43	
6	o	25	19	22	6	o	25	19	36	2	31	56	
7	o	29	32	36	7	o	29	33	37	2	36	10	
8	1	3	45	50	8	o	33	46	38	2	40	23	
9	1	7	59	3	9	o	37	59	39	2	44	36	
10	1	12	12	17	10	o	42	12	40	2	48	49	
11	1	16	25	31	11	o	46	26	41	2	53	3	
12	1	20	38	45	12	o	50	39	42	2	57	16	
13	1	24	51	58	13	o	54	52	43	3	1	29	
14	1	29	5	12	14	o	59	5	44	3	5	42	
15	2	3	18	26	15	1	3	18	45	3	9	55	
16	2	7	31	39	16	1	7	32	46	3	14	9	
17	2	11	44	53	17	1	11	45	47	3	18	22	
18	2	15	58	7	18	1	15	58	48	3	22	35	
19	2	20	11	21	19	1	20	11	49	3	26	48	
20	2	24	24	34	20	1	24	25	50	3	31	2	
21	2	28	37	48	21	1	28	38	51	3	35	15	
22	3	2	51	2	22	1	32	51	52	3	39	28	
23	3	7	4	15	23	1	37	4	53	3	43	41	
24	3	11	17	30	24	1	41	18	54	3	47	55	
					25	1	45	31	55	3	52	8	
					26	1	49	44	56	3	56	21	
					27	1	53	57	57	4	0	34	
					28	1	58	11	58	4	4	48	
					29	2	2	24	59	4	9	1	
					30	2	6	37	60	4	13	14	

In Leap-Year, after February, add the Motion of a Day more to the rest. And the Radix of this Satellite I have taken December 31 at Noon, 1683. Jupiter's Heliocentrick Place was then 5f. 13° 22' 57". Its Half-Stay in the Shadow is 1<sup>h</sup> 25' Mot. 5° 58' 45".



76 A Table of the Motion of the third Satellite of *Jupiter*

Years Cur rent.	Motion.	Years Com.	Motion.	Days.	Motion.
	f. o / "		f. o / "		f. o / "
1681	2 22 15 57	1	11 5 35 40	1	1 20 14 4
1701	6 25 19 36	2	10 11 11 19	2	3 10 28 8
1721	10 28 23 14	3	9 16 46 59	3	5 0 42 13
1722	10 3 38 53	4	10 12 36 44	4	6 20 56 17
1723	9 9 34 33	5	9 18 12 23	5	8 11 10 21
1724	8 15 10 13	6	8 23 48 3	6	10 1 24 25
1725	9 10 59 57	7	7 29 23 42	7	11 21 38 30
1726	8 16 35 37	8	8 25 13 27	8	1 11 52 34
1727	7 22 11 17	9	8 0 49 7	9	3 2 6 38
1728	6 27 46 57	10	7 6 24 46	10	4 22 20 42
1729	7 23 36 41	11	6 12 0 26	11	6 12 34 46
1730	6 29 12 21	12	7 7 50 11	12	8 2 48 51
1731	6 4 48 1	13	6 13 25 50	13	9 23 2 55
1732	5 10 23 40	14	5 19 1 30	14	11 13 16 59
1733	6 6 13 24	15	4 24 37 9	15	1 3 31 3
1734	5 11 49 4	16	5 20 26 54	16	2 23 45 7
1735	4 17 24 44	17	4 26 2 34	17	4 13 59 12
1736	3 23 0 24	18	4 1 38 14	18	6 4 13 16
1737	4 18 50 8	19	3 7 13 53	19	7 24 27 20
1738	3 24 25 48	20	4 3 3 38	20	9 14 41 24
1739	3 0 1 28	40	8 6 7 16	21	11 4 55 29
1740	2 5 37 8	60	0 9 10 54	22	0 25 9 33
1741	3 1 26 52	80	4 12 14 32	23	2 15 23 37
1742	2 7 2 32	100	8 15 18 10	24	4 5 37 41
1743	1 12 38 11	200	5 0 36 21	25	5 25 51 45
1744	0 18 13 51	300	1 15 54 31	26	7 16 5 50
1745	1 4 3 35	400	10 1 12 42	27	9 6 19 54
Jan.	0 0 0 0	500	6 16 30 52	28	10 26 33 58
Feb.	3 27 16 11	600	3 1 49 3	29	0 16 48 2
Mar.	2 23 50 9	700	11 17 7 13	30	2 7 2 7
Apr.	6 21 6 20	800	8 2 25 24	31	3 27 16 11
May.	8 28 8 26	900	4 17 43 34		
June	0 25 24 37	1000	1 3 1 45		
July	3 2 26 43	2000	2 6 3 30		
Aug.	6 29 42 54	3000	3 9 5 15		
Sept.	10 26 59 5	4000	4 12 7 0		
Oct.	1 4 1 12	5000	5 15 8 45		
Nov.	5 1 17 22	6000	6 18 10 30		
Dec.	7 8 19 29	7000	7 21 12 15		



A Table of the Motion of the third Satellite of *Jupiter*. 77

H	Motion.				/	o	/	//	/	o	/	"
	f	o	l	p								
1	o	2	5	35	1	o	2	6	31	1	4	53
2	o	4	11	10	2	o	4	11	32	1	6	59
3	o	6	16	45	3	o	6	17	33	1	9	4
4	o	8	22	21	4	o	8	22	34	1	11	10
5	o	10	27	56	5	o	10	28	35	1	13	15
6	o	12	33	31	6	o	12	33	36	1	15	21
7	o	14	39	6	7	o	14	39	37	1	17	27
8	o	16	44	41	8	o	16	45	38	1	19	32
9	o	18	50	16	9	o	18	50	39	1	21	38
10	o	20	55	52	10	o	20	56	40	1	23	43
11	o	23	1	27	11	o	23	1	41	1	25	49
12	o	23	7	2	12	o	25	7	42	1	27	54
13	o	27	12	37	13	o	27	13	43	1	30	0
14	o	29	18	12	14	o	29	18	44	1	32	6
15	1	1	23	47	15	o	31	24	45	1	34	11
16	1	3	29	23	16	o	33	29	46	1	36	17
17	1	5	34	58	17	o	35	35	47	1	38	22
18	1	7	40	33	18	o	37	40	48	1	40	28
19	1	9	46	8	19	o	39	46	49	1	42	33
20	1	11	51	43	20	o	41	52	50	1	44	39
21	1	13	57	18	21	o	43	57	51	1	46	46
22	1	16	2	54	22	o	46	3	52	1	48	50
23	1	18	8	29	23	o	48	8	53	1	50	57
24	1	20	14	4	24	o	50	14	54	1	53	1
					25	o	52	20	55	1	55	7
					26	o	54	25	56	1	57	14
					27	o	56	31	57	1	59	18
					28	o	58	36	58	2	1	25
					29	1	0	42	59	2	3	29
					30	1	2	47	60	2	5	35

In Leap-Year, after *February*, add the Motion of a Day more to the rest. And the Radix of this Satellite I have taken *December 31* at Noon, 1683. *Jupiter's* Heliocentric Place was then  $5^{\circ} 13' 22'' 57''$ . Half-Stay in the Shadow  $1^{\text{h}} 18'$  Mor.  $2^{\circ} 43' 15''$ .



78 A Table of the Motion of the fourth Satellite of *Jupiter*.

Years Cur- rent	Motion.	Years Com.	Motion.	Days.	Motion.
	f. ° / "		f. ° / "		f. ° / "
1681	8 12 56 51	1	9 13 5 7	1	0 21 29 16
1701	8 22 5 31	2	6 26 10 14	2	1 12 58 33
1721	9 1 14 10	3	4 9 15 21	3	2 4 27 49
1722	6 14 19 17	4	2 13 49 44	4	2 25 57 6
1723	3 27 24 24	5	11 26 54 51	5	3 17 26 22
1724	1 10 29 31	6	9 9 59 58	6	4 8 55 39
1725	11 15 3 54	7	6 23 5 4	7	5 0 24 55
1726	8 28 9 1	8	4 27 39 28	8	5 21 54 12
1727	6 11 14 8	9	2 10 44 35	9	6 13 23 28
1728	3 24 19 14	10	11 23 49 41	10	7 4 52 45
1729	1 28 53 38	11	9 6 54 48	11	7 26 22 1
1730	11 11 58 45	12	7 11 29 12	12	8 17 51 18
1731	8 25 3 51	13	4 24 34 18	13	9 9 20 34
1732	6 8 8 58	14	2 7 39 25	14	10 0 49 51
1733	4 12 43 32	15	11 20 44 32	15	10 22 19 7
1734	1 25 48 28	16	9 25 18 15	16	11 13 48 23
1735	11 8 53 35	17	7 8 24 2	17	0 5 17 40
1736	8 21 58 42	18	4 21 29 9	18	0 26 46 56
1737	6 26 33 5	19	2 4 34 16	19	1 18 16 13
1738	4 9 38 12	20	0 9 8 39	20	2 9 45 29
1739	1 22 43 19	40	0 18 17 19	21	3 1 14 46
1740	11 5 48 26	60	0 27 25 58	22	3 22 44 2
1741	9 10 22 50	80	1 6 34 38	23	4 14 13 19
1742	6 23 27 56	100	1 15 43 17	24	5 5 42 35
1743	4 6 23 3	200	3 1 26 35	25	5 27 11 52
1744	1 19 38 10	300	4 17 9 52	26	6 18 41 8
1745	11 24 12 34	400	6 2 53 9	27	7 10 10 25
Jan.	0 0 0 0	500	7 18 36 27	28	8 1 39 41
Feb.	10 6 7 30	600	9 4 19 44	29	8 23 8 58
Mar.	6 7 47 11	700	10 20 3 1	30	9 14 38 14
Apr.	4 13 54 41	800	0 5 46 19	31	10 6 7 30
May	1 28 32 55	900	1 21 29 36		
June	0 4 40 26	1000	3 7 12 53		
July	9 19 18 40	2000	6 14 25 46		
Aug.	7 25 26 10	3000	9 21 38 39		
Sept.	6 1 33 40	4000	0 28 51 32		
Oct.	3 16 11 54	5000	4 6 4 25		
Nov.	1 22 19 24	6000	7 13 17 18		
Dec.	11 26 57 38	7000	10 20 30 11		



A Table of the Motion of the fourth Satellite of *Jupiter*. 79

H	°	'	"	H	°	'	"
/	/	"	'''	/	/	"	'''
"	''	'''	''''	"	''	'''	''''
1	0	53	43	31	25	45	19
2	1	47	26	32	28	39	2
3	2	41	10	33	29	32	45
4	3	34	53	34	30	26	28
5	4	28	36	35	31	20	11
6	5	22	19	36	32	13	55
7	6	16	2	37	33	7	38
8	7	9	45	38	34	1	21
9	8	3	29	39	34	55	4
10	8	57	12	40	35	48	47
11	9	50	55	41	36	42	31
12	10	44	38	42	37	36	14
13	11	38	21	43	38	29	57
14	12	32	5	44	39	23	40
15	13	25	48	45	40	17	23
16	14	19	31	46	41	11	6
17	15	13	14	47	42	4	50
18	16	6	57	48	42	58	33
19	17	0	40	49	43	52	16
20	17	54	24	50	44	45	59
21	18	48	7	51	45	39	42
22	19	41	50	52	46	33	26
23	20	35	33	53	47	27	9
24	21	29	16	54	48	20	52
25	22	23	0	55	49	14	35
26	23	16	43	56	50	8	18
27	24	10	26	57	51	2	1
28	25	4	9	58	51	55	45
29	25	57	52	59	52	49	28
30	26	51	35	60	53	43	11

In Leap-Year, after *February*, add the Motion of a Day more to the rest. And the Radix of this Satellite I have taken *December 31* at Noon, 1683. *Jupiter's* Heliocentric Place was then  $5^{\circ} 13' 22'' 57''$ . Half-Stay in the Shadow  $1^{\text{h}} 35'$ . Mor.  $1^{\circ} 25' 3''$ .



A Table of the Distances of the Satellites of *Jupiter* from his Body, in Semidiameters and Decimal Parts of *Jupiter's* Globe.

Satellite 1.				
°	0 Conseq. 6 Antec.	1 Conseq. 7 Antec.	2 Conseq. 8 Antec.	°
0	0.	1.8592	3.7183	30
3	0.1859	2.0451	3.9042	27
6	0.3718	2.2310	4.0901	24
9	0.5577	2.4169	4.2760	21
12	0.7436	2.6028	4.4619	18
15	0.9300	2.7887	4.6478	15
18	1.1155	2.9746	4.8337	12
21	1.3014	3.1606	5.0196	9
24	1.4873	3.3465	5.2056	6
27	1.6732	3.5324	5.3915	3
30	1.8592	3.7183	5.5780	0
Satellite 2.				
°	0.	2.9590	5.9180	30
3	0.2959	3.2549	6.2139	27
6	0.5918	3.5508	6.5098	24
9	0.8877	3.8467	6.8057	21
12	1.1836	4.1426	7.1016	18
15	1.4795	4.4385	7.3975	15
18	1.7754	4.7344	7.6934	12
21	2.0713	5.0303	7.9893	9
24	2.3672	5.3262	8.2852	6
27	2.6631	5.6221	8.5811	3
30	2.9590	5.9180	8.8760	0
5 Conseq. 11 Antec.	4 Conseq. 10 Antec.	3 Conseq. 9 Antec.		



A Table of the Distances of the Satellites of *Jupiter* from his Body, in Semidiameters and Decimal Parts of *Jupiter's* Globe.

Satellite 3.				
0	0 Conseq. 6 Antec.	1 Conseq. 7 Antec.	2 Conseq. 8 Antec.	0
0	0.	4.72	9.44	30
3	0.472	5.192	9.912	27
6	0.944	5.664	10.384	24
9	1.416	6.136	10.856	21
12	1.888	6.608	11.328	18
15	2.36	7.08	11.8	15
18	2.832	7.552	12.272	12
21	3.304	8.024	12.744	9
24	3.776	8.496	13.216	6
27	4.248	8.968	13.688	3
30	4.72	9.44	14.159	0
Satellite 4.				
0	0.	8.301	16.602	30
3	0.8301	9.1311	17.4321	27
6	1.6602	9.9612	18.2622	24
9	2.4903	10.7913	19.0923	21
12	3.3204	11.6214	19.9224	18
15	4.1505	12.4515	20.7525	15
18	4.9806	13.2816	21.5826	12
21	5.8107	14.1117	22.4127	9
24	6.6408	14.9418	23.2428	6
27	7.4709	15.7719	24.0729	3
30	8.301	16.602	24.903	0
	5 Conseq. 11 Antec.	4 Conseq. 10 Antec.	3 Conseq. 9 Antec.	



82 A Table of the Motion of the first Satellite of Saturn.

Years Current.	Motion.	Years Com.	Motion.	Days.	Motion.
	f. o ' "		f. o ' "		f. o ' "
1681	9 28 34 39	1	4 4 35 1	1	6 10 41 51
1701	4 23 43 6	2	8 9 10 1	2	0 21 23 42
1721	11 18 52 33	3	0 13 45 2	3	7 2 5 33
1722	3 23 27 34	4	10 29 1 53	4	1 12 47 24
1723	7 28 2 34	5	3 3 36 54	5	7 23 29 15
1724	0 2 37 35	6	7 8 11 55	6	2 4 11 6
1725	10 17 54 27	7	11 12 46 55	7	8 14 52 57
1726	2 22 29 27	8	9 28 3 47	8	2 25 34 48
1727	6 27 4 28	9	2 2 38 47	9	9 6 16 39
1728	11 1 39 28	10	6 7 13 48	10	3 16 58 29
1729	9 16 56 20	11	10 11 48 49	11	9 27 40 20
1730	1 21 31 21	12	8 27 5 40	12	4 8 22 11
1731	5 26 6 21	13	1 1 40 41	13	10 19 4 2
1732	10 0 41 22	14	5 6 15 41	14	4 29 45 53
1733	8 15 58 13	15	9 10 50 42	15	11 10 27 44
1734	0 20 33 14	16	7 26 7 34	16	5 21 9 35
1735	4 25 8 15	17	0 0 42 34	17	0 1 51 26
1736	8 29 43 15	18	4 5 17 35	18	6 12 33 17
1737	7 15 0 7	19	8 9 52 35	19	0 23 15 8
1738	11 19 35 7	20	6 25 9 27	20	7 3 56 59
1739	3 24 10 8	40	1 20 18 55	21	1 14 38 50
1740	7 28 45 9	60	8 15 28 22	22	7 25 20 41
1741	6 14 2 0	80	3 10 37 50	23	2 6 2 32
1742	10 18 37 1	100	10 5 47 17	24	8 16 44 23
1743	2 23 12 1	200	8 11 34 34	25	2 27 16 14
1744	6 27 47 2	300	6 17 21 51	26	9 8 8 5
1745	5 13 3 53	400	4 23 9 8	27	3 18 49 56
Jan.	0 0 0 0	500	2 28 56 25	28	9 29 31 47
Feb.	5 1 37 19	600	1 4 43 42	29	4 10 13 38
Mar.	3 1 9 6	700	11 10 30 59	30	10 20 55 28
Apr.	8 2 46 26	800	9 16 18 16	31	5 1 37 19
May	6 23 41 55	900	7 22 5 33		
June	11 25 19 14	1000	5 27 52 50		
July	10 16 14 53	2000	11 25 45 40		
Aug	3 17 52 13	3000	5 23 38 30		
Sept.	8 19 29 32	4000	11 21 31 20		
Oct.	7 10 24 51	5000	5 19 24 10		
Nov.	0 12 2 11	6000	11 17 17 0		
Dec.	11 2 57 39	7000	5 15 9 50		



H.	Motion.				/	°	'	"	/	°	'	"								
	f.	o	/	"									///	"	///	///	"	///	"	///
1	0	7	56	45	1	0	7	57	31	4	6	19								
2	0	15	53	29	2	0	15	53	32	4	14	15								
3	0	23	50	14	3	0	23	50	33	4	22	12								
4	1	1	46	58	4	0	31	47	34	4	30	9								
5	1	9	43	43	5	0	39	44	35	4	38	6								
6	1	17	40	28	6	0	47	40	36	4	46	2								
7	1	25	37	12	7	0	55	37	37	4	53	59								
8	2	3	33	57	8	1	3	34	38	5	1	56								
9	2	11	30	42	9	1	11	31	39	5	9	53								
10	2	19	27	26	10	1	19	27	40	5	17	49								
11	2	27	24	11	11	1	27	24	41	5	25	46								
12	3	5	20	55	12	1	35	21	42	5	33	43								
13	3	13	17	40	13	1	43	18	43	5	41	40								
14	3	21	14	25	14	1	51	14	44	5	49	36								
15	3	29	11	9	15	1	59	11	45	5	57	33								
16	4	7	7	54	16	2	7	8	46	6	5	30								
17	4	15	4	38	17	2	15	4	47	6	13	26								
18	4	23	1	23	18	2	23	1	48	6	21	23								
19	5	0	58	8	19	2	30	58	49	6	29	20								
20	5	8	54	52	20	2	38	55	50	6	37	17								
21	5	16	51	37	21	2	46	51	51	6	45	13								
22	5	24	48	22	22	2	54	48	52	6	53	10								
23	6	2	45	7	23	3	2	45	53	7	1	7								
24	6	10	41	51	24	3	10	42	54	7	9	4								
					25	3	18	38	55	7	17	0								
					26	3	26	35	56	7	24	57								
					27	3	34	32	57	7	32	54								
					28	3	42	29	58	7	40	51								
					29	3	50	25	59	7	48	57								
					30	3	58	22	60	7	56	45								

In Leap-Year, after February, add the Motion of a Day more to the rest.

Its Nodes are in  $\text{M} \times 21^\circ$ . Inclination  $31^\circ$ .

The Epoche of this Satellite I have taken, Anno 1713, December 31 at Noon,  $\text{h}$  Geocentric Place was then  $5 \text{ f. } 11^\circ 6' 9''$ .



84 A Table of the Motion of the second Satellite of Saturn.

Years Cur- rent.	Motion.	Years Com.	Motion.	Days.	Motion.
	f. ° / "		f. ° / "		f. ° / "
1681	11 5 28 41	1	4 10 2 52	1	4 11 32 4
1701	11 24 6 19	2	8 20 5 44	2	8 23 4 7
1721	0 12 43 57	3	1 0 8 36	3	1 4 36 11
1722	4 22 46 49	4	9 21 43 32	4	5 16 8 15
1723	9 2 49 41	5	2 1 46 23	5	9 27 40 19
1724	1 12 52 33	6	6 11 49 15	6	2 9 12 22
1725	10 4 27 29	7	10 21 52 7	7	6 20 44 26
1726	2 14 30 21	8	7 13 27 3	8	11 2 16 30
1727	6 24 33 13	9	11 23 29 55	9	3 13 48 34
1728	11 4 36 5	10	4 3 32 47	10	7 25 20 37
1729	7 26 11 0	11	8 13 39 39	11	0 6 52 41
1730	0 6 13 52	12	5 5 10 35	12	4 18 24 45
1731	4 16 16 44	13	9 15 13 27	13	8 29 56 49
1732	8 26 19 36	14	1 25 16 19	14	1 11 28 52
1733	5 17 54 32	15	6 5 19 10	15	5 23 0 56
1734	9 27 57 24	16	2 26 54 6	16	10 4 33 0
1735	2 8 0 16	17	7 6 56 58	17	2 16 5 4
1736	6 18 3 8	18	11 16 59 50	18	6 27 37 7
1737	3 9 38 3	19	3 27 2 42	19	11 9 9 11
1738	7 19 40 55	20	0 18 37 38	20	3 20 41 15
1739	11 29 43 47	40	1 7 15 16	21	8 2 13 19
1740	4 9 46 39	60	1 25 52 54	22	0 13 45 22
1741	1 1 21 35	80	2 14 30 31	23	4 25 17 26
1742	5 11 24 27	100	3 3 8 9	24	9 6 49 30
1743	9 21 27 18	200	6 6 16 19	25	1 18 21 34
1744	2 1 30 10	300	9 9 24 28	26	5 29 53 37
1745	10 23 5 7	400	0 12 32 37	27	10 11 25 41
Jan.	0 0 0 0	500	3 15 40 47	28	2 22 57 45
Feb.	3 27 33 57	600	6 18 48 56	29	7 4 29 49
Mar.	6 20 31 42	700	9 21 57 5	30	11 16 1 52
Apr.	10 18 5 39	800	0 25 5 15	31	3 27 33 57
May.	10 4 7 31	900	3 28 13 24		
June	2 1 41 28	1000	7 1 21 33		
July	1 17 43 20	2000	2 2 43 6		
Aug.	5 15 17 17	3000	9 4 4 39		
Sept.	9 12 51 14	4000	4 5 26 12		
Oct.	8 28 53 6	5000	11 6 47 45		
Nov.	0 26 27 3	6000	6 8 9 18		
Dec.	0 12 28 55	7000	1 9 30 51		



A Table of the Motion of the second Satellite of Saturn. 85

H	Motion.				/	o	/	//	/	o	/	"
	f.	q	l	"								
1	0	5	28	50	1	0	5	29	31	2	49	55
2	0	10	57	40	2	0	10	58	32	2	55	24
3	0	16	26	30	3	0	16	26	33	3	0	52
4	0	21	55	21	4	0	21	55	34	3	6	20
5	0	27	24	11	5	0	27	24	35	3	11	49
6	1	2	53	1	6	0	32	53	36	3	17	18
7	1	8	21	51	7	0	38	22	37	3	22	47
8	1	13	50	41	8	0	43	51	38	3	28	16
9	1	19	19	31	9	0	49	19	39	3	33	44
10	1	24	48	21	10	0	54	48	40	3	39	14
11	2	0	17	12	11	1	0	17	41	3	44	43
12	2	5	46	2	12	1	5	46	42	3	50	12
13	2	11	14	52	13	1	11	15	43	3	55	41
14	2	16	43	42	14	1	16	44	44	4	1	9
15	2	22	12	32	15	1	22	12	45	4	6	36
16	2	27	41	22	16	1	27	41	46	4	12	5
17	3	3	10	13	17	1	33	10	47	4	17	34
18	3	8	39	3	18	1	38	39	48	4	23	3
19	3	14	7	53	19	1	44	8	49	4	28	32
20	3	19	36	43	20	1	49	38	50	4	34	1
21	3	25	5	33	21	1	55	7	51	4	39	29
22	4	0	34	23	22	2	0	35	52	4	44	58
23	4	6	3	13	23	2	6	4	53	4	50	27
24	4	11	32	4	24	2	11	33	54	4	55	56
					25	2	17	2	55	5	1	25
					26	2	22	31	56	5	6	54
					27	2	28	0	57	5	12	22
					28	2	33	28	58	5	17	51
					29	2	38	57	59	5	23	20
					30	2	44	26	60	5	28	50

In Leap-Year, after February, add the Motion of a Day more to the rest.

Its Nodes are in  $\text{III} \times 21^\circ$ . Inclination  $31^\circ$ .

The Epoche of this Satellite I have taken, Anno 1713, December 31 at Noon,  $\text{h}$  Geocentric Place was then  $51^\circ 11' 9''$ .



86 A Table of the Motion of the third Satellite of *Saturn*.

Years Cur- rent	Motion.				Years Com.	Motion.				Days.	Motion.			
	f.	o	'	"		f.	o	'	"		f.	o	'	"
1681	5	4	23	10	1	9	17	2	0	1	2	19	41	26
1701	5	23	30	23	2	7	4	4	0	2	5	9	22	52
1721	6	12	37	36	3	4	21	6	0	3	7	29	4	17
1722	3	29	39	36	4	4	27	49	26	4	10	18	45	43
1723	1	16	41	36	5	2	14	51	27	5	1	8	27	9
1724	11	3	43	36	6	0	1	53	27	6	3	28	8	35
1725	11	10	27	2	7	9	18	55	27	7	6	17	50	1
1726	8	27	29	2	8	9	25	38	53	8	9	7	31	26
1727	6	14	32	3	9	7	12	40	53	9	11	27	12	52
1728	4	1	33	3	10	4	29	42	53	10	2	16	54	18
1729	4	8	16	29	11	2	16	44	53	11	5	6	35	44
1730	1	25	18	29	12	2	23	28	19	12	7	26	17	10
1731	11	12	20	29	13	0	10	30	20	13	10	15	58	35
1732	8	29	22	29	14	9	27	32	20	14	1	5	40	1
1733	9	6	5	55	15	7	14	34	20	15	3	25	21	27
1734	6	23	7	55	16	7	21	17	46	16	6	15	2	53
1735	4	10	9	55	17	5	8	19	46	17	9	4	44	19
1736	1	27	11	56	18	2	25	21	46	18	11	24	25	44
1737	2	3	55	22	19	0	12	23	46	19	2	14	7	10
1738	11	20	57	22	20	0	19	7	13	20	5	3	48	36
1739	9	7	59	22	40	1	8	14	25	21	7	23	30	2
1740	6	25	1	22	60	1	27	21	38	22	10	13	11	28
1741	7	1	44	48	80	2	16	28	51	23	1	2	52	53
1742	4	18	46	48	100	3	5	36	3	24	3	22	34	19
1743	2	5	48	48	200	6	11	12	7	25	6	12	15	45
1744	11	22	50	49	300	9	16	48	10	26	9	1	57	11
1745	11	29	34	15	400	0	22	24	13	27	11	21	38	37
Jan.	0	0	0	0	500	3	28	0	17	28	2	11	20	2
Feb.	10	10	24	20	600	7	3	36	20	29	5	1	1	28
Mar.	0	21	44	22	700	10	9	12	23	30	7	20	42	54
Apr.	11	2	8	43	800	1	14	48	27	31	10	10	24	20
May	6	22	51	37	900	4	20	24	30					
June	5	3	15	57	1000	7	26	0	23					
July	0	23	58	51	2000	3	22	1	6					
Aug.	11	4	23	11	3000	7	14	2	12					
Sept.	9	14	47	31	4000	3	10	2	45					
Oct.	5	5	30	25	5000	11	6	3	18					
Nov.	3	15	54	45	6000	7	2	3	51					
Dec.	11	6	37	39	7000	2	28	4	24					



A Table of the Motion of the third Satellite of *Saturn*. 87

H.	Motion.				/	o	/	"	/	o	/	"
	f.	o	l	"								
1	o	3	19	14	1	o	3	19	31	1	42	56
2	o	6	38	27	2	o	6	38	32	1	46	15
3	o	9	57	41	3	o	9	58	33	1	49	34
4	o	13	16	54	4	o	13	17	34	1	52	53
5	o	16	36	8	5	o	16	36	35	1	56	13
6	o	19	55	21	6	o	19	55	36	1	59	32
7	o	23	14	35	7	o	23	15	37	2	2	51
8	o	26	33	49	8	o	26	34	38	2	6	10
9	o	29	53	2	9	o	29	53	39	2	9	29
10	1	3	12	16	10	o	33	12	40	2	12	49
11	1	6	31	29	11	o	36	31	41	2	16	8
12	1	9	50	43	12	o	39	51	42	2	19	27
13	1	13	9	56	13	o	43	10	43	2	22	46
14	1	16	29	10	14	o	46	29	44	2	26	6
15	1	19	48	23	15	o	49	48	45	2	29	25
16	1	23	7	37	16	o	53	7	46	2	32	44
17	1	26	26	51	17	o	56	27	47	2	36	3
18	1	29	46	4	18	o	59	46	48	2	39	22
19	2	3	5	18	19	1	3	5	49	2	42	42
20	2	6	24	31	20	1	6	24	50	2	46	1
21	2	9	43	45	21	1	9	44	51	2	49	20
22	2	13	2	58	22	1	13	3	52	2	52	39
23	2	16	22	12	23	1	16	22	53	2	55	58
24	2	19	41	26	24	1	19	41	54	2	59	18
					25	1	23	o	55	3	2	37
					26	1	26	20	56	3	5	56
					27	1	29	39	57	3	9	15
					28	1	32	58	58	3	12	35
					29	1	36	17	59	3	15	54
					30	1	39	36	60	3	19	14

In Leap-Year, after *February*, add the Motion of a Day more to the rest.

Its Nodes are in ♃ ♋ 21°. Inclination 31.

The Epoche of this Satellite I have taken, *Anno* 1713, *December* 31 at Noon, ♁ Geocentric Place was then 5f. 11° 6' 9".



88 A Table of the Motion of the fourth Satellite of Saturn.

Years Current.	Motion.	Years Com.	Motion.	Days.	Motion.
	f. o / "		f. o / "		f. o / "
1681	1 0 48 56	1	10 20 35 19	1	0 22 34 37
1701	2 15 28 19	2	9 11 10 37	2	1 15 9 14
1721	4 0 7 42	3	8 1 45 56	3	2 7 43 51
1722	2 20 43 1	4	7 14 55 52	4	3 0 18 28
1723	1 11 18 20	5	6 5 31 11	5	3 22 53 5
1724	0 1 53 39	6	4 26 6 30	6	4 15 27 42
1725	11 15 3 35	7	3 16 41 49	7	5 8 2 19
1726	10 5 38 54	8	2 29 51 45	8	6 0 36 56
1727	8 26 14 13	9	1 20 27 4	9	6 23 11 33
1728	7 16 49 31	10	0 11 2 23	10	7 15 46 10
1729	6 29 59 27	11	11 1 37 42	11	8 8 20 47
1730	5 20 34 46	12	10 14 47 38	12	9 0 55 24
1731	4 11 10 5	13	9 5 22 57	13	9 23 30 1
1732	3 1 45 24	14	7 25 58 15	14	10 16 4 38
1733	2 14 55 20	15	6 16 33 34	15	11 8 39 15
1734	1 5 30 39	16	5 29 43 30	16	0 1 13 52
1735	11 26 5 58	17	4 20 18 49	17	0 23 48 29
1736	10 16 41 17	18	3 10 54 8	18	1 16 23 6
1737	10 29 51 13	19	2 1 29 27	19	2 8 57 43
1738	9 20 26 31	20	1 14 39 23	20	3 1 32 20
1739	8 11 1 50	40	2 29 18 46	21	3 24 6 57
1740	7 1 37 9	60	4 13 58 9	22	4 16 41 34
1741	5 14 47 5	80	5 28 37 32	23	5 9 16 11
1742	4 5 22 24	100	7 13 16 55	24	6 1 50 48
1743	2 25 57 43	200	2 26 33 50	25	6 24 25 25
1744	1 16 33 2	300	10 9 50 44	26	7 17 0 2
1745	0 29 42 58	400	5 23 7 39	27	8 9 34 39
Jan.	0 0 0 0	500	1 6 24 34	28	9 2 9 16
Feb.	11 9 53 8	600	8 19 44 29	29	9 24 43 53
Mar.	8 12 2 26	700	4 2 58 24	30	10 17 18 32
Apr.	7 21 55 34	800	11 16 15 19	31	11 9 53 8
May	6 9 14 6	900	6 29 32 13		
June	5 19 7 14	1000	2 12 49 8		
July	4 6 25 40	2000	4 25 38 16		
Aug.	3 16 18 54	3000	7 8 27 24		
Sept.	2 26 12 2	4000	9 21 16 32		
Oct.	1 13 30 34	5000	0 4 5 40		
Nov.	0 23 23 42	6000	2 16 54 48		
Dec.	11 10 42 14	7000	4 29 43 56		



A Table of the Motion of the fourth Satellite of Saturn. 89

H.	°	'	"	H.	°	'	"
1	0	56	27	31	29	9	43
2	1	52	53	32	30	6	9
3	2	49	20	33	31	2	36
4	3	45	46	34	31	59	2
5	4	42	13	35	32	55	29
6	5	38	39	36	33	51	55
7	6	35	6	37	34	48	22
8	7	31	32	38	35	44	48
9	8	27	59	39	36	41	15
10	9	24	25	40	37	37	41
11	10	20	52	41	38	34	8
12	11	17	18	42	39	30	34
13	12	13	45	43	40	27	1
14	13	10	11	44	41	23	27
15	14	6	38	45	42	19	54
16	15	3	5	46	43	16	21
17	15	59	31	47	44	12	47
18	16	55	58	48	45	9	14
19	17	52	24	49	46	5	40
20	18	48	51	50	47	2	7
21	19	45	17	51	47	58	53
22	20	41	44	52	48	55	0
23	21	38	10	53	49	51	26
24	22	34	37	54	50	47	53
25	23	31	3	55	51	44	19
26	24	27	30	56	52	40	46
27	25	23	56	57	53	37	12
28	26	20	23	58	54	33	39
29	27	16	49	59	55	30	5
30	28	13	16	60	56	26	32

In Leap-Year, after February, add the Motion of a Day more to the rest.

Its Nodes are in  $\text{♊}$   $21^\circ$ . Inclination  $31^\circ$ .

The Epoche of this Satellite I have taken, Anno 1713, December 31 at Noon.  $\text{h}$  Geocentric Place was then  $5 \text{ s. } 11^\circ 6' 9''$ .



90 A Table of the Motion of the fifth Satellite of Saturn.

Years Current	Motion.	Years Com.	Motion.	Days	Motion.
	f. o ' "		f. o ' "		f. o ' "
1681	0 9 20 39	1	7 6 30 20	1	0 4 32 18
1701	1 12 8 47	2	2 13 0 40	2	0 9 4 36
1721	2 14 56 55	3	9 19 31 0	3	0 13 36 54
1722	9 21 27 15	4	5 0 33 38	4	0 18 9 13
1723	4 27 57 35	5	0 7 3 58	5	0 22 41 31
1724	0 4 27 55	6	7 13 34 17	6	0 27 13 49
1725	7 15 30 33	7	2 20 4 37	7	1 1 46 7
1726	2 22 0 53	8	10 1 7 15	8	1 6 18 25
1727	9 28 31 13	9	5 7 37 35	9	1 10 50 43
1728	5 5 1 33	10	0 14 7 55	10	1 15 23 1
1729	0 16 4 11	11	7 20 38 15	11	1 19 55 19
1730	7 22 34 31	12	3 1 40 53	12	1 24 27 38
1731	2 29 4 51	13	10 8 11 13	13	1 28 59 56
1732	10 5 35 11	14	5 14 41 33	14	2 3 32 14
1733	5 16 37 49	15	0 21 11 53	15	2 8 4 32
1734	0 23 8 8	16	8 2 14 31	16	2 12 36 50
1735	7 29 38 28	17	3 8 44 51	17	2 17 9 8
1736	3 6 8 48	18	10 15 15 10	18	2 21 41 26
1737	10 17 11 26	19	5 21 45 30	19	2 26 13 45
1738	5 23 41 46	20	1 2 48 8	20	3 0 46 3
1739	1 0 12 6	40	2 5 36 17	21	3 5 18 21
1740	8 6 42 26	60	3 8 24 25	22	3 9 50 39
1741	3 17 45 4	80	4 11 12 34	23	3 14 22 57
1742	10 24 15 24	100	5 14 0 42	24	3 18 55 15
1743	6 0 45 44	200	10 28 1 24	25	3 23 27 33
1744	1 7 16 4	300	4 12 2 6	26	3 27 59 51
1745	8 18 18 42	400	9 26 2 48	27	4 2 32 10
Jan.	0 0 0 0	500	3 10 3 30	28	4 7 4 28
Feb.	4 20 41 22	600	8 24 4 12	29	4 11 36 46
Mar.	8 27 45 50	700	2 8 4 55	30	4 16 9 4
Apr.	1 18 27 12	800	7 22 5 37	31	4 20 41 22
May	6 4 36 16	900	1 6 6 19		
June	10 25 17 38	1000	6 20 7 1		
July	3 11 26 42	2000	1 10 14 2		
Aug.	8 2 8 5	3000	8 0 21 3		
Sept.	0 22 42 27	4000	2 20 28 4		
Oct.	5 8 58 31	5000	9 10 35 5		
Nov.	9 29 39 53	6000	4 0 42 6		
Dec.	2 15 48 57	7000	10 20 49 7		



A Table of the Motion of the fifth Satellite of *Saturn*. 91

H.	°	'	"	H.	°	'	"
/	/	"	'''	/	/	"	'''
//	//	'''	''''	//	//	'''	''''
1	0	11	21	31	5	51	43
2	0	22	41	32	6	3	4
3	0	34	2	33	6	14	25
4	0	45	23	34	6	25	45
5	0	56	44	35	6	37	6
6	1	8	4	36	6	48	27
7	1	19	25	37	6	59	48
8	1	30	46	38	7	11	8
9	1	42	7	39	7	22	29
10	1	53	27	40	7	33	50
11	2	4	48	41	7	45	11
12	2	16	9	42	7	56	31
13	2	27	30	43	8	7	52
14	2	38	50	44	8	19	13
15	2	50	11	45	8	30	34
16	3	1	32	46	8	41	54
17	3	12	53	47	8	53	15
18	3	24	13	48	9	4	36
19	3	35	34	49	9	15	57
20	3	46	55	50	9	27	17
21	3	58	16	51	9	38	38
22	4	9	36	52	9	49	59
23	4	20	57	53	10	1	20
24	4	32	18	54	10	12	40
25	4	43	39	55	10	24	1
26	4	54	59	56	10	35	22
27	5	6	20	57	10	46	43
28	5	17	41	58	10	57	3
29	5	29	2	59	11	9	24
30	5	40	22	60	11	20	45

In Leap-Year, after *February*, add the Motion of a Day more to the rest.

Its Nodes are in  $\text{III} \times 4^\circ$ . Inclination  $16^\circ$ .

The Epoche of this Satellite I have taken, *Anno* 1713, December 31 at Noon.  $\text{h}$  Geocentric Place was then  $55^\circ 11' 6'' 9''$ .



A Table of the Distances of *Saturn's* Satellites, from his Body in Semidiameters and Decimal Parts of his Globe.

Satellite 1.				Satellite 2.				Satellite 3.			
0	Con. 1	Con. 2	Con.	0	Con. 1	Con. 2	Con.	0	Con. 1	Con. 2	Con.
6	Ant. 7	Ant. 8	Ant.	6	Ant. 7	Ant. 8	Ant.	6	Ant. 7	Ant. 8	Ant.
0	0.0000	1.446	2.892	30	0.0000	1.833	3.706	30	0.0000	2.5881	5.1762
3	0.1446	1.5906	3.0366	27	0.1853	2.0383	3.8913	27	3.02588	2.847	5.435
6	0.2892	1.7352	3.1812	24	0.3706	2.2236	4.0766	24	6.05176	3.1057	5.6938
9	0.4338	1.8798	3.3258	21	0.5559	2.4089	4.2619	21	9.07764	3.3645	5.9526
12	0.5784	2.0244	3.4704	18	0.7412	2.5942	4.4472	18	12.10352	3.6233	6.2114
15	0.723	2.169	3.615	15	0.9265	2.7795	4.6325	15	15.12940	3.8821	6.4702
18	0.8676	2.3136	3.7596	12	1.1118	2.9648	4.8178	12	18.15529	4.141	6.7291
21	1.0122	2.4582	3.9042	9	1.2971	3.1501	5.0031	9	21.18117	4.3998	6.9879
24	1.1568	2.6028	4.0488	6	1.4824	3.3354	5.1184	6	24.20705	4.6586	7.2467
27	1.3014	2.7474	4.1934	3	1.6677	3.5207	5.3787	3	27.23293	4.9174	7.5055
30	1.446	2.892	4.34	0	1.853	3.706	5.5593	0	30.25881	4.1762	7.7643
5	Con. 4	Con. 3	Con.	5	Con. 4	Con. 3	Con.	5	Con. 4	Con. 3	Con.
11	Ant. 10	Ant. 9	Ant.	11	Ant. 10	Ant. 9	Ant.	11	Ant. 10	Ant. 9	Ant.



A Table of the Distances of the Satellites of Saturn from his Body, in Semidiameters and Decimal Parts of his Globe.

Satellite 4.				
0	0 Conseq. 6 Antec.	1 Conseq. 7 Antec.	2 Conseq. 8 Antec.	0
0	0.0000	6.	12.	30
3	0.6	6.6	12.6	27
6	1.2	7.2	13.2	24
9	1.8	7.8	13.8	21
12	2.4	8.4	14.4	18
15	3.	9.	15.	15
18	3.6	9.6	15.6	12
21	4.2	10.2	16.2	9
24	4.8	10.8	16.8	6
27	5.4	11.4	17.4	3
30	6.	12.	18.	0

Satellite 5.				
0	0.0000	17.4859	34.9719	30
3	1.7486	19.2345	36.7205	27
6	3.4972	20.9831	38.469	24
9	5.2458	22.7317	40.2176	21
12	6.9944	24.4803	41.9662	18
15	8.743	26.2289	43.7148	15
18	10.4916	27.9775	45.4634	12
21	12.2402	29.7261	47.212	9
24	13.9887	31.4747	48.9606	6
27	15.7373	33.2233	50.7092	3
30	17.4859	34.9719	52.4578	0

5 Conseq. 11 Antec.	4 Conseq. 10 Antec.	3 Conseq. 9 Antec.



A Table of the Number of Days from the first of  
*January* to any Day in the Year.

Days.	Jan.	Feb.	Mar.	Apr.	May	June	July	Aug.	Sept.	Oct.	Nov.	Dec.
1	1	32	60	91	121	152	182	213	244	274	305	335
2	2	33	61	92	122	153	183	214	245	275	306	336
3	3	34	62	93	123	154	184	215	246	276	307	337
4	4	35	63	94	124	155	185	216	247	277	308	338
5	5	36	64	95	125	156	186	217	248	278	309	339
6	6	37	65	96	126	157	187	218	249	279	310	340
7	7	38	66	97	127	158	188	219	250	280	311	341
8	8	39	67	98	128	159	189	220	251	281	312	342
9	9	40	68	99	129	160	190	221	252	282	313	343
10	10	41	69	100	130	161	191	222	253	283	314	344
11	11	42	70	101	131	162	192	223	254	284	315	345
12	12	43	71	102	132	163	193	224	255	285	316	346
13	13	44	72	103	133	164	194	225	256	286	317	347
14	14	45	73	104	134	165	195	226	257	287	318	348
15	15	46	74	105	135	166	196	227	258	288	319	349
16	16	47	75	106	136	167	197	228	259	289	320	350
17	17	48	76	107	137	168	198	229	260	290	321	351
18	18	49	77	108	138	169	199	230	261	291	322	352
19	19	50	78	109	139	170	200	231	262	292	323	353
20	20	51	79	110	140	171	201	232	263	293	324	354
21	21	52	80	111	141	172	202	233	264	294	325	355
22	22	53	81	112	142	173	203	234	265	295	326	356
23	23	54	82	113	143	174	204	235	266	296	327	357
24	24	55	83	114	144	175	205	236	267	297	328	358
25	25	56	84	115	145	176	206	237	268	298	329	359
26	26	57	85	116	146	177	207	238	269	299	330	360
27	27	58	86	117	147	178	208	239	270	300	331	361
28	28	59	87	118	148	179	209	240	271	301	332	362
29	29	—	88	119	149	180	210	241	272	302	333	363
30	30	—	89	120	150	181	211	242	273	303	334	364
31	31	—	90	—	151	—	212	243	—	304	—	365



**I**N Page 70, I have given you a Table of the Hourly Motions, Apparent Semidiameter, and Horizontal Parallaxes of the Sun and Moon; but you are to observe, that that Table is computed to the middle State of the Lunar Orbit, not having regard to the Change of her Eccentricity: But that you may have these things true at all times, observe the following Method.

To find the Horizontal Parallax and Apparent Semidiameter of the Moon, according to the Theory.

*Example.* To the Place of the Moon Jan. 1, 1729, I would know her Horizontal Parallax, and Apparent Semidiameter, which are obtain'd from Figure 2. as follows

*Operation.*

As *f.*  $\triangle$  E L F, the Ellip. Equat.  $3^{\circ} 4' 38''$  Co-Ar. 1.270174  
 To E F the Double Eccentricity 90794 - - - - 4.958057  
 So *f.*  $\triangle$  L F E, Mean Anomaly 37 51 52 - - - 9.787897  
 To E L, Dist.  $\Delta$  from the Earth 1037834 - - - 6.016128

Now say,

As present Dist. $\Delta$ à $\odot$	— — — —	6.016128
To her mean Distance	— — — —	6.000000
So <i>f.</i> mean Horizontal Parallax $\Delta$	57' 30" - -	8.223357
To <i>f.</i> present Horizontal Parallax	55 24 - -	8.207229

2. For the Apparent Semidiameter  $\Delta$  at the same time, say by the Logistical Logarithms,

As the mean Horizont. Parall. 57' 30" LL Co-Ar.	815
To mean Semidiameter — 15 45 — —	5809
So is the present Horiz. Paral. 55 24 — — —	346
<hr/>	
To the present Apparent Semid. 15 11 — —	5970

And after the same manner may you find the Moon's Horizontal Parallax to the second Example, January 2. at Noon 1729, to be 56 14", and the Apparent Semidiameter 15' 24".

For



For the true hourly Motions of the Sun and Moon in Eclipses, &c. calculate their Places to half an Hour before, and to half an Hour after the equal Time of the true Conjunction or Opposition, and so you will gain their true hourly Motions at that time.

N. B. The Table in Page 94, is to be used when you find the Day of the Retrogradation of any Planet by the 11th Precent, of my *Compleat System of Astronomy*.

FINIS



POLITECHNIKA GDAŃSKA  
Z ZASOBÓW  
POLITECHNIKI GDANSKIEJ  
N 500029



